

1 **A parameterization of convective dust storms**
2 **for models with mass-flux convection schemes**

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ABSTRACT

15 Cold pool outflows, generated by downdrafts from moist convection, can
16 generate strong winds and therefore uplift of mineral dust. These so-called
17 “haboob” convective dust storms occur over all major dust source areas world-
18 wide and contribute substantially to emissions in northern Africa, the world’s
19 largest source. Most large-scale models lack convective dust storms, because
20 they do not resolve moist convection, relying instead on convection schemes.
21 We suggest a parameterization of convective dust storms to account for their
22 contribution in such large-scale models. The parameterization is based on a
23 simple conceptual model, in which the downdraft mass flux from the convec-
24 tion scheme spreads out radially in a cylindrical cold pool. The parameteri-
25 zation is tested with a set of Unified Model runs for June and July 2006 over
26 West Africa. It is calibrated with a convection-permitting run, and applied to
27 a convection-parameterized run. The parameterization successfully produces
28 the extensive area of dust-generating winds from cold pool outflows over the
29 southern Sahara. However, this area extends farther to the east and dust gen-
30 erating winds occur earlier in the day than in the convection-permitting run.
31 These biases are due to biases in the convection scheme. It is found that
32 the location and timing of dust-generating winds are weakly sensitive to the
33 parameters of the conceptual model. The results demonstrate that a simple
34 parameterization has the potential to correct a major and long-standing limi-
35 tation in global dust models.

36 **1. Introduction**

37 In a thunderstorm, the melting, evaporation, and sublimation of hydrometeors generate down-
38 drafts that form a spreading cold pool at low levels (Byers 1949). The cold pool is denser than its
39 environment and therefore spreads as a density current (e.g., Simpson 1999). The cold pool plays
40 a dual role in the life cycle of the thunderstorm: it increases the low-level atmospheric stability
41 and locally inhibits convection, but additionally lifts the surrounding, warmer air and triggers new
42 convective cells (Byers 1949).

43 The cold pool outflow creates a front of wind gusts at its leading edge. Over arid grounds,
44 the wind gusts can be strong enough to lift mineral dust. This process was first documented
45 in peer-reviewed literature for Karthoum and described as “haboob” (Sutton 1925). Since then,
46 haboobs have been reported over all major sources of mineral dust worldwide (see Knippertz
47 2014, and references therein). Dust uplift is found in cold pool outflows of different space and
48 time scales: mesoscale convective systems (Houze 2004) can produce long-lived haboobs (Roberts
49 and Knippertz 2014); small, strong downdrafts (microbursts, Fujita 1985) can produce short-lived
50 haboobs (Miller et al. 2008); even small cold pools from precipitating congestus can produce dust
51 uplift (Marsham et al. 2009). As all processes are related to convection, they are referred to as
52 convective dust storms.

53 Convective dust storms of different origins have been observed over the Sahara during recent
54 field campaigns: created by orographic convection over the northwestern Sahara (during SA-
55 MUM, Knippertz et al. 2007); embedded within the monsoon flow over the southern Sahara (dur-
56 ing AMMA, Flamant et al. 2007; Bou Karam et al. 2008) and over the western Sahara (during
57 GERBILS, Marsham et al. 2008b); and over the central Sahara, from locally generated moist con-
58 vection, as well as mesoscale convective systems that propagate from the Sahel (during FENNEC,

59 Marsham et al. 2013b; Allen et al. 2013). Observational (Marsham et al. 2008b, 2013b) and mod-
60 eling studies (Heinold et al. 2013) suggest that convective dust storms contribute a large fraction
61 of dust emission over the Sahara in summer. The Sahara is the main source of mineral dust world-
62 wide, and convective dust storms may contribute to the local and remote impacts of Saharan dust
63 on health, oceanic biochemistry, and atmospheric dynamics (see Knippertz and Todd 2012, for a
64 review of mineral dust over the Sahara).

65 Investigating the systematic impact of convective dust storms is challenging: the ground ob-
66 servation network is sparse over the Sahara, and convective clouds often hide dust in satellite
67 observations (Heinold et al. 2013; Kocha et al. 2013). Furthermore, most operational models lack
68 convective dust storms (Marsham et al. 2011; Garcia-Carreras et al. 2013), since they do not ex-
69 plicitly resolve convection and rely on parameterization schemes. Parameterization schemes lack
70 microbursts, because they do not account for subgrid-scale winds. Parameterization schemes also
71 lack mesoscale convective systems, because they do not account for grid-scale organization of con-
72 vection (e.g., Knippertz and Todd 2012). A parameterization of convective dust storms is needed
73 to account for their contribution to dust uplift in large-scale models.

74 Several authors have parameterized wind gusts according to convective downdrafts: Nakamura
75 et al. (1996) assumed conservation of horizontal momentum in downdrafts to compute peak wind
76 gusts in numerical weather prediction models; Redelsperger et al. (2000) defined subgrid gustiness
77 as a function of the downdraft mass flux to enhance surface fluxes in global circulation models;
78 Cakmur et al. (2004) scaled a probability distribution of subgrid wind with the downdraft mass
79 flux to compute dust uplift in global circulation models. Building on these previous studies, we
80 suggest a parameterization of subgrid winds for dust uplift based on the downdraft mass flux of a
81 convective parameterization scheme. Our parameterization aims at remaining simple, in order to
82 be applied online or offline to any model with mass-flux convection scheme. It contrasts with the

83 integrated approach of Hourdin et al. (2014), which improves the representation of wind and dust
84 emissions in a global model – although it does not address the issue of convective dust storms – but
85 requires a complete modification of subgrid parameterization schemes. Our parameterization also
86 complements statistical downscaling methods, which improve dust emissions in global models but
87 still lack the contribution from convective dust storms, such as the one by Ridley et al. (2013).

88 Section 2 describes the configuration of the model runs used to formulate the parameterization,
89 compares their representation of cold pools and dust generating winds, and details the reference
90 used to calibrate the parameterization. Section 3 explains and illustrates the conceptual model of
91 the parameterization and its tuning. Section 4 gives the results of the parameterization for both the
92 geographical distribution and diurnal cycle. Finally, Section 5 concludes the paper and discusses
93 perspectives for future work.

94 **2. Model runs**

95 *a. Configuration*

96 The parameterization of convective dust storms is based on a set of model runs with the UK
97 Met Office Unified Model. The Unified Model uses a seamless approach, from weather forecast
98 to climate projection and from limited area to global domain (Walters et al. 2011). In the frame-
99 work of the Cascade project, the model was run in a limited area configuration over West Africa
100 at different spatial resolutions, with and without parameterisations of moist convection, and for
101 different time periods during the summer 2006. The Cascade project allowed an investigation of
102 the representation of tropical convection (Pearson et al. 2010, 2014; Birch et al. 2014a), its im-
103 pact on the monsoon (Marsham et al. 2013a; Birch et al. 2014b), and its impact on dust emission
104 (Marsham et al. 2011; Heinold et al. 2013).

105 The present study is mainly based on two runs with 4-km and 12-km grid spacings for the 60-
106 day period 1 June to 30 July 2006. Diagnostics for convective mass fluxes, which are essential for
107 the formulation of the parameterization of convective dust storms, were saved out during this time
108 period only. Additional runs for the 10-day period 25 July to 3 August 2006 are also discussed,
109 because the model was run at higher resolution with 1.5-km grid spacing for this time period, in
110 addition to the 4-km and 12-km grid spacings. As convective mass fluxes were not saved out for
111 this 10-day period, the additional runs cannot be used for the parameterization of convective dust
112 storms. The relevant characteristics of the different runs are summarized in Table 1.

113 The model was run over limited area domains on a rotated cylindrical grid. Figure 1 illustrates
114 the orography, soil fraction, and surface roughness over the 12-km domain. Figure 1a further dis-
115 plays the 4-km and 1.5-km domains. Operational analyses from the European Centre for Medium-
116 Range Weather Forecasts (ECMWF) provided the initial conditions and lateral boundaries for the
117 12-km runs (Table 1). The 12-km runs provided the lateral boundaries conditions for the nested
118 4-km runs. The 4-km run for the 10-day period in turn provided the lateral boundaries for the
119 nested 1.5-km run. Terrain-following hybrid coordinates were used in the vertical, with 70 levels
120 starting at 2.5 m in the 4-km and 1.5-km runs, and with 38 levels starting at 10 m in the 12-km run
121 (Table 1). The model configuration is detailed in Pearson et al. (2010).

122 The 1.5-km and 4-km runs fundamentally differ in their representation of convection as com-
123 pared to the 12-km run: the convection is permitted to develop explicitly with 1.5-km and 4-
124 km grid spacings, while it is parameterized with 12-km grid spacing (Table 1). In the Unified
125 Model, the parameterization of moist convection is based on a convective available potential en-
126 ergy (CAPE) closure (Gregory and Rowntree 1990). Following a parcel theory modified by en-
127 trainment and detrainment, an ensemble of subgrid convective clouds is described by updraft and
128 downdraft mass fluxes. Updrafts are initiated if a layer is positively buoyant; ascent occurs until

129 the parcel becomes negatively buoyant. In turn, downdrafts are initiated as a fraction of updrafts
130 if a layer is negatively buoyant; descent occurs until the parcel becomes positively buoyant or too
131 close to the surface.

132 *b. Representation of cold pools*

133 Figure 2 compares the representation of cold pools in the 1.5-km, 4-km, and 12-km runs on
134 31 July 2006 (10-day period, Table 1). The respective peak of the diurnal cycle of precipitation
135 is illustrated; it occurs at 12 UTC in the 12-km run (Fig. 2g) instead of 17 UTC in the 1.5-km
136 and 4-km runs (Figs. 2a,d). The parameterization scheme triggers convection too early in the 12-
137 km run (Marsham et al. 2013a; Birch et al. 2014b; Pearson et al. 2014), which is a common and
138 well-documented issue in tropical regions (Yang and Slingo 2001; Dai 2006; Nikulin et al. 2012;
139 Bechtold et al. 2014). Note that the 3 runs are not expected to look the same at any particular
140 time, because they are only constrained at the lateral boundaries. The panels in Fig. 2 are used for
141 illustration purposes only.

142 In both the 1.5-km and 4-km run, convective cells produce strong precipitation above 10 mm h^{-1}
143 (Figs. 2a,d). The evaporation, melting, and sublimation of hydrometeors create cold pools at low
144 levels with temperature contrast above 5 K (Figs. 2b,e). The outflow of cold pools produces strong
145 surface winds above 10 m s^{-1} (Figs. 2c,f). Convective cells produce small, circular cold pools,
146 which grow and merge into larger, more complex structures. In contrast, the convection scheme
147 produces weak precipitation below 10 mm h^{-1} in the 12-km run (Fig. 2g). The evaporation of
148 precipitation is too weak and too widespread to produce distinct cold pools (Fig. 2h). The 12-km
149 run therefore lacks high winds resulting from convective cold pool outflows (Fig. 2i).

150 This qualitative comparison suggests that the 4-km and 1.5-km runs offer a similar representa-
151 tion of convection and strongly contrast with the 12-km run. Earlier studies showed that convection

152 in the 1.5-km and 4-km runs occurs with a good timing compared to satellite observations, while
153 convection occurs too early in the 12-km run (Marsham et al. 2013a; Birch et al. 2014b; Pearson
154 et al. 2014). Furthermore, the development and growth of convective organization is weakly sen-
155 sitive to the resolution between the 1.5-km and 4-km runs (Pearson et al. 2014). Weisman et al.
156 (1997) also found that the structure and evolution of mesoscale convective systems varied little
157 between runs with 4-km and 1-km grid spacing, although convection was slightly delayed with the
158 coarser grid spacing. In contrast with the 1.5-km and 4-km runs, the 12-km run lacks organized
159 convection (Birch et al. 2014a; Pearson et al. 2014) and cold pools (Marsham et al. 2011, 2013a;
160 Heinold et al. 2013).

161 A quantitative comparison is given by the frequency of surface wind speed over the Sahara in the
162 runs during the 10-day period (Fig. 3). While the 12-km run distribution drops near 12 m s^{-1} , the
163 4-km run matches the 1.5-km run and captures the tail of distribution up to 20 m s^{-1} . Convective
164 dust storms contribute most of the tail of distribution (not shown). This further supports that the
165 representation of cold pool outflows is similar in the 4-km and 1.5-km runs. Johnson et al. (2014)
166 also show that the timing and structure of a convective outflow are successfully represented with a
167 4-km grid spacing. The 4-km run is then the only available run that explicitly represents convection
168 and captures the cold pool outflows during the 60-day period, for which the convective mass flux
169 diagnostics were saved out (Table 1). As observations are sparse over the Sahara, the 4-km run is
170 used as a reference for the parameterization of convective dust storms. It provides robust statistics
171 with a large number (many hundreds) of convective dust storms that develop during the 60-day
172 period.

173 *c. Dust uplift potential*

174 Dust uplift occurs when the friction velocity reaches a threshold that depends on soil properties
175 such as mineralogy, roughness elements, and moisture (Marticorena and Bergametti 1995; Shao
176 and Lu 2000). The friction velocity was not saved out in the runs. We therefore estimate dust
177 uplift from the 10-m wind speed, which largely controls the friction velocity. Several authors have
178 directly computed the friction velocity from the 10-m wind speed (e.g., Cakmur et al. 2004; Miller
179 et al. 2008; Ridley et al. 2013; Fiedler et al. 2013). Here we follow Marsham et al. (2011) and
180 compute the dust uplift potential

$$\text{DUP} = \nu U_{10}^3 \left(1 + \frac{U_t}{U_{10}} \right) \left(1 - \frac{U_t^2}{U_{10}^2} \right), \quad (1)$$

181 with ν the fraction of bare soil, U_{10} the 10-m wind speed, and $U_t = 7 \text{ m s}^{-1}$ a fixed threshold
182 for dust uplift. The DUP isolates the atmospheric control from the soil control on dust uplift, and
183 thus can easily be computed offline without a full model for dust emission. Heinold et al. (2013)
184 showed that DUP is largely consistent with both the diurnal cycle and the geographical distribution
185 of dust emission fluxes from such a full model. Marsham et al. (2013b) further showed that DUP
186 correlates with observed dust over the central Sahara.

187 The geographical distribution of DUP exhibits similar patterns in the 4-km and 12-km runs
188 (Fig. 4). Highest DUP is found over the Saharan heat low region from eastern Mauritania to
189 northern Mali (18-22°N, 12-2°W) and over the Bodélé Depression in northern Chad (16-20°N,
190 15-20°E). High DUP is found over southwestern Algeria (24-27°N, 5-0°W), where it is related to
191 the flow around the Hoggar Mountains (Birch et al. 2012), and over northeastern Niger (20-24°N,
192 10-18°E). High DUP is also found along the coast of Mauritania and Western Sahara, where the
193 Atlantic inflow produces strong winds during the afternoon and evening (Grams et al. 2010).

194 Apart from the Atlantic coast, the areas of high DUP coincide with the areas of highest fraction
195 of bare soil (Fig. 1b). However, the pattern of bare soil does not directly impact the pattern of DUP:
196 omitting ν in Eq. 1 produces a similar pattern of DUP (not shown). Instead, the low roughness
197 length over bare soil (Fig. 1c) allows for strong winds that result in high DUP (Fig. 4). The sharp
198 border in DUP along the Sahel (near 16°N in Fig. 4) matches the strong gradient in roughness
199 length (Fig. 1c). The roughness length increases over mountain ranges, because it accounts for
200 subgrid orography (Fig. 1a). High roughness length prevents strong winds and DUP over the
201 Tibesti ($19\text{-}24^{\circ}\text{N}$, $16\text{-}20^{\circ}\text{E}$) and Hoggar ($22\text{-}27^{\circ}\text{N}$, $3\text{-}13^{\circ}\text{E}$) mountain ranges (Fig. 4).

202 Figure 5 displays the diurnal cycle of DUP over the Sahara. A strong peak occurs in the morning
203 and is attributed to the breakdown of the nocturnal low-level jet (Knippertz 2008; Fiedler et al.
204 2013). The 12-km run underestimates the amplitude of the peak compared to the 4-km run (Fig. 5).
205 In contrast, the 12-km run overestimated the amplitude of the peak during the 10-day period, due
206 to a deeper Saharan heat low, and thus a stronger pressure gradient compared to the 4-km run
207 (Marsham et al. 2013a; Heinold et al. 2013). Here, the 12-km run exhibits a shallower Saharan
208 heat low than the 4-km run (contours in Fig. 4). This demonstrates how sensitive the monsoon
209 circulation is to the time period and representation of convection in a given model (Marsham et al.
210 2013a). The weaker pressure gradient in the 12-km run results in weaker nocturnal low-level jets
211 and therefore weaker DUP in the morning compared to the 4-km run (Fig. 5). Heinold et al. (2013)
212 showed that low-level jets can form in aged cold pools, such that some of the differences between
213 the two runs may indirectly be related to the lack of organized convection in the 12-km run.

214 A second, weaker peak in DUP occurs in the afternoon, in both 4-km and 12-km runs (Fig. 5).
215 This peak is attributed to dry convection in the boundary layer, which reaches its peak in the
216 afternoon and which was observed to enhance dust uplift (Chaboureau et al. 2007; Marsham et al.
217 2008a). DUP then remains high in the evening in the 4-km run, while it drops in the 12-km run.

218 The weaker DUP in the 12-km run was attributed to the lack of convective dust storms in the
219 evening during the 10-day period (Marsham et al. 2011; Heinold et al. 2013). The contribution of
220 convective dust storms to DUP in the 4-km run is discussed below.

221 *d. Identification of convective dust storms*

222 Convective dust storms need to be identified in the 4-km run, which is used as a reference
223 to calibrate the parameterization. Following Heinold et al. (2013), surface winds are attributed
224 to convective dust storms if they occur within 40 km of a grid point of rapid cooling and strong
225 vertical velocities. These conditions are met at the leading edge of cold pool outflows (see example
226 of cold pool outflow in Section 3a). Additional conditions in potential temperature and wind
227 divergence suggested by Heinold et al. (2013) were found redundant here with the conditions in
228 cooling and vertical velocity, respectively.

229 A visual inspection of several cold pool outflows in the 4-km run delivered thresholds $\dot{T}_t = -1$ K
230 h^{-1} for temperature tendency and $|w|_t = 0.5$ m s^{-1} for vertical velocity of up- and downdrafts. The
231 1-h temperature tendency is computed on the 133-m model level and defined as the anomaly with
232 respect to the 5-day average of the diurnal cycle, while the vertical velocity is taken on the 1605-m
233 model level. The choice of 1-h tendency and 5-day average was constrained by the organization of
234 model data, while the choice of model levels was driven by the strongest signature of cold pools
235 in temperature tendency and vertical velocity.

236 The thresholds are close to those defined by Heinold et al. (2013). Figure 6 shows the diurnal
237 cycle of identified convective dust storms using a range of thresholds \dot{T}_t for temperature tendency
238 and $|w|_t$ for vertical velocity. Regardless of thresholds, DUP from convective dust storms quickly
239 increases from 13 UTC to reach its peak at 18 UTC, consistent with the peak rain at this time
240 (Marsham et al. 2013a; Birch et al. 2014b; Pearson et al. 2014). This contributes to the overall

241 DUP peak in the afternoon (blue curve in Fig. 5). DUP from convective dust storms then declines
242 until 06 UTC (Fig. 6), when rainfall is low and the strong surface stable layer inhibits cold pool
243 momentum from reaching the surface. A weak peak occurs at 09 UTC, during the breakdown
244 of the nocturnal low-level jet (Fig. 5). This is consistent with cold pool momentum being mixed
245 down to the surface as dry convection erodes the stable layer (Heinold et al. 2013).

246 Heinold et al. (2013) found low sensitivity to the exact thresholds used. Here, multiplying \dot{T}_t or
247 $|w|_t$ by a factor of 2 increases DUP by 33% and 24%, respectively (red curves in Fig. 6). Dividing
248 \dot{T}_t or $|w|_t$ by a factor of 2 decreases DUP by 42% and 20%, respectively (blue curves in Fig. 6).
249 These results suggest that the uncertainty in the contribution of convective dust storms is on the
250 order of 30%. The uncertainty accounts both for spurious rejection of cold pool outflows and for
251 spurious identification of other processes. While isolated cold pools are distinct, however, their
252 identification is ambiguous when they are embedded in the monsoon flow or evolve into nocturnal
253 low-level jets (Heinold et al. 2013).

254 3. Conceptual model

255 In order to address the problem of lacking cold pool dust emission in models with parameterized
256 convection, we now present the conceptual model on the basis of which our parameterization of
257 convective dust storms is built. Section 3a presents the general formulation, while Section 3b
258 shows an illustrative example, which is used to tune the parameterization in Section 3c.

259 a. Formulation

260 The parameterization is based on the conceptual model of convective dust storms that is illus-
261 trated in Fig. 7: the downdraft mass flux M_{dd} (in kg s^{-1}) spreads out radially in a cylindrical cold
262 pool of radius R and height h . To ensure conservation of mass, the propagation speed of the cold

263 pool must be

$$C = \frac{M_{dd}}{2\pi R h \rho}, \quad (2)$$

264 with ρ the average density of the cold pool. The conceptual model matches a developing cold pool
265 in the 4-km run: a strong convective downdraft (Figs. 8b,d) spreads out radially in a cylindrical
266 cold pool and creates strong winds at its leading edge (Figs. 8a,c).

267 When a cold pool propagates as a density current, its radius increases and its propagation speed
268 decreases¹. In contrast, convective parameterizations assume the quasi equilibrium of subgrid
269 boundary-layer processes (Bechtold et al. 2014). A parameterization of propagating, subgrid cold
270 pools therefore requires the complete coupling with the parameterization of subgrid convection
271 (Grandpeix and Lafore 2010). Such a coupling is beyond the scope of our work. We rather
272 base our parameterization on a single, static cold pool of representative size (with sensitivity to
273 assumptions of size tested in Section 3c). The conceptual model is therefore independent of the
274 model time step if applied online, or of the temporal sampling of model output if applied offline.

275 Surface friction lifts the leading edge of a density current, which forms a “nose” (Simpson 1999).
276 The developing cold pool in the 4-km run exhibits such a nose with strongest wind at height $z_{max} \approx$
277 100 m (Fig. 8c). Below z_{max} , turbulence mixes the surface layer. We assume the surface layer has
278 constant potential temperature, i.e., neutral stability, and that below z_{max} the radial wind speed
279 follows a logarithmic profile

$$U_r(z) = \frac{u^*}{\kappa} \ln(z_{max}/z_0), \quad (3)$$

280 with z the height above ground, u^* the friction velocity, $\kappa = 0.41$ the von Karman constant, and
281 z_0 the roughness length. Above z_{max} , the radial wind speed decreases with height (Fig. 8c). While
282 the internal flow of the cold pool is directed forward at low levels, it is directed backward closer to

¹The theoretical propagation speed of a cold pool follows $C \propto R^{-1/3} \propto t^{-1/4}$ if the downdraft mass flux is sustained (Parker 1996), and $C \propto R^{-1} \propto t^{-1/2}$ if the downdraft mass flux is stopped at some point (Simpson 1999).

283 the top levels (Simpson 1999). For simplicity, we assume the radial wind speed decreases linearly
 284 with height above z_{max} and vanishes at height h . The thin black arrows illustrate the vertical profile
 285 of the radial wind in Fig. 7.

286 Combining the logarithmic profile below z_{max} and the linear profile above, the maximum radial
 287 wind speed at the leading edge must satisfy,

$$U_r(z_{max}) = \alpha C \quad (4)$$

288 at height z_{max} , with

$$\alpha = h \left(z_{max} \frac{\ln(z_{max}/z_0) - 1}{\ln(z_{max}/z_0)} + \frac{1}{2}(h - z_{max}) \right)^{-1} \quad (5)$$

289 to ensure conservation of mass. With typical values $z_{max} = 100$ m and $z_0 = 10^{-3}$ m, α increases
 290 from $\alpha \approx 1.1$ for $h = z_{max}$ to $\alpha = 2$ for $h \gg z_{max}$; a height $h = 240$ m delivers the value $\alpha = 1.5$
 291 that was observed in thunderstorm outflows (Goff 1976).

292 Within the cold pool, we assume M_{dd} to be homogeneous. To ensure conservation of mass, the
 293 radial wind speed must read

$$U_r(r) = \frac{r}{R} U_r(R), \quad (6)$$

294 with r the distance from the center of the cold pool (thin black arrows in Fig. 7). Based on
 295 observations of strong downdrafts, Holmes and Oliver (2000) also used Eq. 6 to describe the wind
 296 speed for $r < R$. In addition, they suggested an empirical model of the form

$$U_r(r) = e^{-\left(\frac{r-R}{R_0}\right)^2} U_r(R) \quad (7)$$

297 for $r > R$, with $R_0 \approx 0.5R$ a radial length scale. We apply this empirical model to account for the
 298 smooth decrease in wind speed beyond the leading edge of the cold pool (Figs. 8a,c).

299 The developing cold pool in the 4-km run exhibits asymmetric wind speeds (Figs. 8a,c), because
 300 the downdraft transports horizontal momentum from higher levels (Figs. 8b,d). Following Parker

301 (1996), we write the steering speed of the cold pool,

$$C_{st} = 0.65U_{env}, \quad (8)$$

302 where U_{env} is the environmental steering wind. The relevant layer for U_{env} is where the downdraft
303 originates from, and not where it spreads out (Fig. 7). We assume that the steering wind within
304 the cold pool (gray arrows) follows the vertical profile of the radial wind (black arrows). The
305 maximum steering wind therefore reads

$$U_{st}(z_{max}) = \alpha C_{st} \quad (9)$$

306 at height z_{max} , with α given by Eq. 5.

307 Following Holmes and Oliver (2000), the total wind is obtained from the vector addition of
308 radial and steering wind,

$$U_{tot}(\mathbf{r}) = \frac{\mathbf{r}}{R}U_r(r) + \mathbf{U}_{st}. \quad (10)$$

309 The conceptual model does not explicitly account for the vertical wind shear. The wind shear
310 sustains cold pools in organized convective systems (Rotunno et al. 1988) but does not impact the
311 propagation of a cold pool as a density current (Parker 1996).

312 *b. Illustration*

313 Equations 2 to 10 describe the conceptual model. In the following, we apply them to the devel-
314 oping cold pool in the 4-km run (Fig. 8). The downdraft mass flux is computed from the vertical
315 velocity w_{dd} of downdrafts as

$$M_{dd} = \int_A \rho w_{dd} dA, \quad (11)$$

316 with A the area of the cold pool. The downdraft mass flux reaches its peak $M_{dd} = 1.5 \times 10^9 \text{ kg s}^{-1}$
317 on the 1605-m model level (Fig. 8b). The average environmental wind within the cold pool reaches

318 $U_{env} = 4.5 \text{ m s}^{-1}$ and blows west-southwestward on the same model level. A visual estimate gives
319 parameters $R = 20 \text{ km}$, $R_0 = 0.33R$ (Fig. 8a), $h = 2 \text{ km}$, and $z_{max} = 100 \text{ m}$ (Fig. 8c); additional
320 parameters are $\rho = 1 \text{ kg m}^{-3}$ and $z_0 = 5 \times 10^{-3} \text{ m}$ in the model run.

321 Given the estimated parameters, the conceptual model yields $C = 6.0 \text{ m s}^{-1}$ (Eq. 2), $\alpha = 1.9$
322 (Eq. 5), $U_r(z_{max}) = 11.5 \text{ m s}^{-1}$ (Eq. 4), and $U_{st}(z_{max}) = 5.7 \text{ m s}^{-1}$ (Eqs. 8 and 9). The radial wind
323 at height z_{max} is computed from Eqs. 6 and 7, then the total wind at height z_{max} is computed from
324 Eq. 10. Finally, the total wind is extrapolated to $z = 10 \text{ m}$ from Eq. 3. Alternatively, the friction
325 velocity can be computed from the total wind in Eq. 3. The total wind is set to vanish at distance
326 $r = R + R_0$ from the center, to avoid the environmental wind extending outside of the cold pool.

327 Figure 9 illustrates the resulting wind field. The conceptual model captures the asymmetric
328 structure of the cold pool outflow and its magnitude in the 4-km run (Figs. 8a,c). The exact
329 intensity of surface winds can be obtained by tuning the parameters carefully. The strong wind
330 speed along the downdraft at the center of the cold pool (Figs. 8c,d) is lacking in the conceptual
331 model (Fig. 9b), but it does not affect the surface wind. New up- and downdrafts at the leading
332 edge of the cold pool (Figs. 8b,d) are also lacking as expected in the conceptual model, but they
333 play a minor role during the early development of the cold pool.

334 The 4-km run exhibits variability in the structure of cold pool outflows (Fig. 2f). The concep-
335 tual model does not account for fine-scale processes that impact the development of cold pools
336 (e.g., surface inhomogeneities, Lothon et al. 2011). However, the crescent shape of surface winds
337 (Fig. 9) matches the typical structure of cold pool outflows in the 4-km run (Fig. 2f). This suggests
338 that the simple assumptions of the conceptual model (Fig. 7) deliver a realistic, albeit idealized,
339 representation of cold pools outflows.

340 *c. Tuning*

341 The downdraft mass flux computed from the vertical velocity of downdrafts (Eq. 11) reaches
342 $M_{dd} = 1.5 \times 10^9 \text{ kg s}^{-1}$ in the developing cold pool of the 4-km run (Fig. 8). In contrast, the down-
343 draft mass flux diagnostic computed in the convective parameterization scheme barely reaches 1.5
344 $\times 10^7 \text{ kg s}^{-1}$ over the Sahara in the 12-km run (Fig. 10a). Two reasons explain this difference in
345 magnitude. Firstly, the radius of parameterized convective cells in the 12-km run must be on the
346 order of one kilometer to remain of subgrid size, while the radius of the developing cold pool in
347 the 4-km run reaches $R = 20 \text{ km}$. Secondly, the downdraft mass flux of the convection scheme is
348 typically too weak, due to the lack of explicit representation of subgrid variability. In particular, a
349 more intense downdraft mass flux would over-stabilize the lower layers (Ben Shipway 2014, UK
350 Met Office, personal communication). Cakmur et al. (2004) scaled the downdraft mass flux of the
351 convection scheme with an empirical constant $\beta = 10$ to compute subgrid wind for dust uplift.
352 Following Cakmur et al. (2004), we scale M_{dd} with an arbitrary factor $f = 10$ in the conceptual
353 model, unless stated otherwise.

354 Several parameters control the wind speed in the conceptual model: the radius of cold pools R
355 (Eq. 2), the height of cold pools h (Eqs. 2 and 5), the height of maximum winds z_{max} (Eqs. 3 and
356 5), and the radial length scale R_0 (Eq. 7). We constrain the geometry of cold pool outflows to
357 reduce the number of free parameters to one: based on the developing cold pool in the 4-km run
358 (Fig. 8), we set $h/R = 0.1$, $z_{max} = 100 \text{ m}$, and $R_0/R = 0.33$. The parameterization now depends
359 on R only. Using a different constraint on the geometry of the cold pool requires a different tuning
360 of R but weakly impacts the resulting DUP.

361 The free parameter R is tuned for the average parameterized DUP to match the average reference
362 DUP (the calibration area is discussed in Section 4). The parameterized DUP is computed from the

363 parameterized subgrid wind and averaged over the grid cells in the 12-km run, while the reference
364 DUP is computed from the model wind attributed to convective dust storms in the 4-km run. Using
365 a trial-and-error method, the best match of the parameterized DUP with the reference DUP is found
366 for a radius of cold pools $R = 2.0$ km. The constraint on the geometry of cold pools gives a height
367 $h = 0.2$ km. Parameterized downdrafts of subgrid scale spread out in cold pools of subgrid scale
368 as expected. Their radius corresponds to the typical radius of microbursts (Fujita 1985).

369 An additional, hidden parameter of the conceptual model is the height at which the environmen-
370 tal wind U_{env} is taken. Figure 10b illustrates the distribution of U_{env} over the Sahara at different
371 model levels in the 12-km run. The distribution of U_{env} is computed where M_{dd} is positive only,
372 i.e. where the parameterization will be applied. Increasing the height between 2210 and 4210 m
373 quickly shifts the distribution to stronger U_{env} . The distribution is more stable below and above
374 this range of heights (not shown). This shows that the chosen level strongly impacts the value
375 of U_{env} in the parameterization. However, the chosen level weakly impacts the surface wind: a
376 typical $U_{env} = 5 \text{ m s}^{-1}$ (Fig. 10b) yields a steering speed of the cold pool $C_{st} = 3 \text{ m s}^{-1}$ (Eq. 8).
377 In comparison, a typical $M_{dd} = 5 \times 10^6 \text{ kg s}^{-1}$ (Fig. 10a) scaled by $f = 10$ yields a propagation
378 speed of the cold pool $C = 20 \text{ m s}^{-1}$ (Eq. 2). The height at which U_{env} is taken is therefore not
379 expected to strongly affect the DUP overall, but may impact DUP locally if high U_{env} combines
380 with low M_{dd} . Here the 3130-m level was chosen as a compromise between weaker and stronger
381 environmental winds (Fig. 10b).

382 4. Space and time distribution of convective dust storms

383 The DUP from convective dust storms is first discussed in the 4-km run. Identified convective
384 dust storms produce DUP over the southern Sahara mainly (around 18°N , Fig. 11a), where the
385 monsoon flow brings the necessary moisture to trigger convection. Highest DUP is found over

386 the Saharan heat low from eastern Mauritania to northern Mali, as for the total DUP (Fig. 4a).
387 In contrast, low DUP is found over the Bodélé Depression in northern Chad and over southwest-
388 ern Algeria (Fig. 11a), consistent with known wind sources that are not related to cold pools in
389 these regions (Washington and Todd 2005; Birch et al. 2012). Local concentrations of DUP are
390 found over southern Algeria and northeastern Niger, in the vicinity of mountain ranges (Fig. 4a),
391 consistent with orographic triggering of moist convection.

392 The parameterization of convective dust storms in the 12-km run succeeds at producing high
393 DUP over the southern Sahara (around 18°N, Fig. 11b). The parameterized DUP is shifted east-
394 ward compared to the DUP from identified convective dust storms in the 4-km run (Fig. 11a). The
395 eastward shift in the location of DUP is due to the eastward shift in the location of precipitation be-
396 tween the 12-km and the 4-km runs (contours in Fig. 11). The location of precipitation is coupled
397 with the pressure gradient of the Saharan heat low (contours in Fig. 4) through the dynamics of the
398 monsoon (Marshall et al. 2013a; Birch et al. 2014b). The parameterized DUP further lacks local
399 concentrations in the vicinity of mountain ranges compared to the 4-km run (Fig. 11), because of
400 the relative lack of moist convection in the vicinity of mountain ranges in the 12-km run.

401 Although most of DUP over the Sahel south of 16°N is attributed to convective dust storms
402 in the 4-km run, it remains small compared to DUP over the Sahara (Figs. 4a and 11a). The
403 parameterized DUP extends farther south across the Sahel (Fig. 11). This appears more realistic
404 than the sharp border in the 4-km run, as convective dust storms have been observed along a
405 transect around 14°N at the beginning of the monsoon season (Marticorena et al. 2010). The high
406 roughness length over the Sahel (Fig. 1c) prevents strong winds in the model runs; it is possibly
407 too high for the beginning of the monsoon season, when the vegetation has not yet developed.

408 High DUP is also attributed to convective dust storms along the coast in the 4-km run (Fig. 11a).
409 The Atlantic inflow is identified as a cold pool outflow, because its front propagates as a density

410 current (Grams et al. 2010). However, the Atlantic inflow does not result from convection; it is
411 therefore excluded from the calibration area (boxes in Fig. 11). The northern and eastern margins
412 of the nested 4-km domain are also excluded from the calibration area to avoid contamination
413 from the lateral boundaries. The calibration area also excludes the area south of 15°N, because the
414 reference 4-km run may underestimate DUP over the Sahel.

415 As seen in Fig. 6, DUP from convective dust storms exhibits a strong diurnal cycle in the 4-
416 km run (Fig. 12, blue curve). Convective dust storms contribute 27 % of the total DUP from 13
417 UTC to 06 UTC and 16 % of the total daily DUP, over the calibration area displayed in Fig. 11a.
418 The parameterized DUP succeeds at exhibiting a strong diurnal cycle (Fig. 12, red curve). As
419 expected, however, the peak of parameterized DUP occurs at 12 UTC instead of 18 UTC in the
420 4-km run, because convection is triggered too early in the 12-km run (Marsham et al. 2013a; Birch
421 et al. 2014b; Pearson et al. 2014). The parameterized DUP then decreases too quickly after the
422 peak since the moist convection is too short-lived in the 12-km run. As the parameterization is
423 calibrated with the daily DUP, the amplitude of the peak is overestimated compared to the 4-km
424 run (Fig. 12). Therefore, the main biases in timing and amplitude of DUP are due to biases in the
425 convective parameterization scheme, and not to the parameterization of convective dust storms.

426 **5. Conclusion**

427 We suggest a parameterization of convective dust storms for models with mass-flux convection
428 schemes. The parameterization is based on a set of Unified Model runs over West Africa for June
429 and July 2006. It is applied to a convection-parameterized run with 12-km grid spacing, which
430 lacks convective dust storms. A convection-permitting run with 4-km grid spacing captures the
431 dynamics of convective dust storms and is used as a reference for validation and tuning.

432 Our conceptual model of convective dust storms follows simple assumptions (Fig. 7). The down-
433 draft mass flux – a known value from the convective parameterization scheme – spreads out radi-
434 ally in a static, cylindrical cold pool. The resulting radial wind adds to the steering wind of the
435 downdraft. Together, they follow a logarithmic profile below the “nose” of the cold pool, and de-
436 crease linearly with height above. The conceptual model reproduces the structure and magnitude
437 of wind speed for a developing cold pool in the reference run.

438 The parameterization produces a distribution of subgrid wind in each grid cell of the 12-km run.
439 It is calibrated to match the integrated dust generating winds (dust uplift potential, DUP) from
440 identified convective dust storms over the Sahara in the reference run. The geometry of the cold
441 pools is constrained in the parameterization, based on a developing cold pool in the reference run.
442 The only free parameter is the radius of the cold pools, which is taken as constant for the whole
443 domain and the whole period. The calibration delivers a radius of 2.0 km, consistent with the
444 subgrid downdraft mass fluxes producing subgrid cold pools.

445 The parameterization of convective dust storms successfully produces high DUP over the south-
446 ern Sahara. The parameterized DUP is more spread out than in the reference run: it lacks local
447 concentrations over the central Sahara and extends farther east over the southern Sahara. Over the
448 Sahel, the parameterized DUP extends farther south and appears more realistic than the reference
449 run, which shows a sharp border at 16°N. The parameterization of convective dust storms also
450 successfully produces a strong diurnal cycle of DUP. The parameterized DUP peaks 6-h earlier
451 and reaches higher amplitude than in the reference run.

452 Compared to the reference run, differences in the geographical distribution of parameterized
453 convective dust storms originate from differences in the monsoon flow between the model runs.
454 Differences in the timing of convective dust storms also originate from differences in the timing of
455 convection between the model runs. The dynamics of the West African monsoon (e.g., Marsham

456 et al. 2013a) and the diurnal cycle of tropical convection (e.g., Bechtold et al. 2014) are know
457 issues for modeling and are topics of active research. These issues are separate from the lack of
458 convective dust storms addressed here and solving them is beyond the scope of this paper.

459 The results suggest that the new parameterization allows a useful estimate of dust uplift due to
460 convective dust storms. The distribution and timing of DUP are weakly sensitive to the parameters
461 of the conceptual model, if the radius of cold pools is carefully calibrated. The main uncertainty
462 originates from the calibration, which is sensitive to the model resolution, the chosen domain and
463 period, the identification of convective dust storms, and the estimate of dust uplift in the reference
464 run. The uncertainty, however, remains small compared to large uncertainties in the estimation of
465 dust uplift from models and observations (Huneeus et al. 2011).

466 As the parameterization produces a distribution of subgrid wind, it can be implemented in a full
467 model for dust emission. If required, the parameterization can alternatively produce a distribution
468 of subgrid friction velocity. A more accurate estimate of dust uplift can then be used instead of
469 the simple DUP to tune the parameterization for the full model. The uplifted dust will then be
470 transported beyond the grid cell, mixed, or deposited by the meteorology of the model. Through
471 both wetting of the soil and scavenging, convective precipitation within a column may reduce the
472 efficiency of convective dust storms in that column in a full dust model. To account for the spatial
473 separation between the gust front and the precipitation in a realistic convective dust storm, the
474 best approach may be to switch off the soil moisture effect and the scavenging during time steps
475 when the parameterization is activated. A more detailed investigation of this effect is left for future
476 applications in a fully online coupled system.

477 Further work is needed to test the sensitivity of the parameterization to different periods, grid
478 spacings, and models. Current parameters of the conceptual model may vary: e.g., the radius of
479 cold pools, which is expected to increase with increasing grid spacing. Parameterized convective

480 dust storms would have more realistic dimensions with grid spacings on the order of 100 km.
481 Additional parameters may be included in the conceptual model: e.g., the vertical wind shear,
482 which is crucial for the organization of convection (Rotunno et al. 1988). If proven robust, the
483 parameterization will substantially improve the representation of a key ingredient to dust emission
484 and allow studies of the impact of convective dust storms in large-scale weather and climate models
485 that use mass-flux convection schemes.

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TABLE 1. Relevant characteristics of the model runs discussed in the text.

Period	Dates	Grid spacing	Vertical levels	Lateral boundaries	Convection	Mass flux diagnostics
10 day	25 July to 3 August	1.5 km	70	4-km run	explicit	
10 day	25 July to 3 August	4 km	70	12-km run	explicit	
10 day	25 July to 3 August	12 km	38	ECMWF analyses	parameterized	not available
60 day	1 June to 30 July	4 km	70	12-km run	explicit	
60 day	1 June to 30 July	12 km	38	ECMWF analyses	parameterized	available

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664 model orography. The displayed area is the northern part of the 4-km domain (Fig. 1). The
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677 of maximum wind. Thin black and gray arrows illustrate the radial and the steering wind
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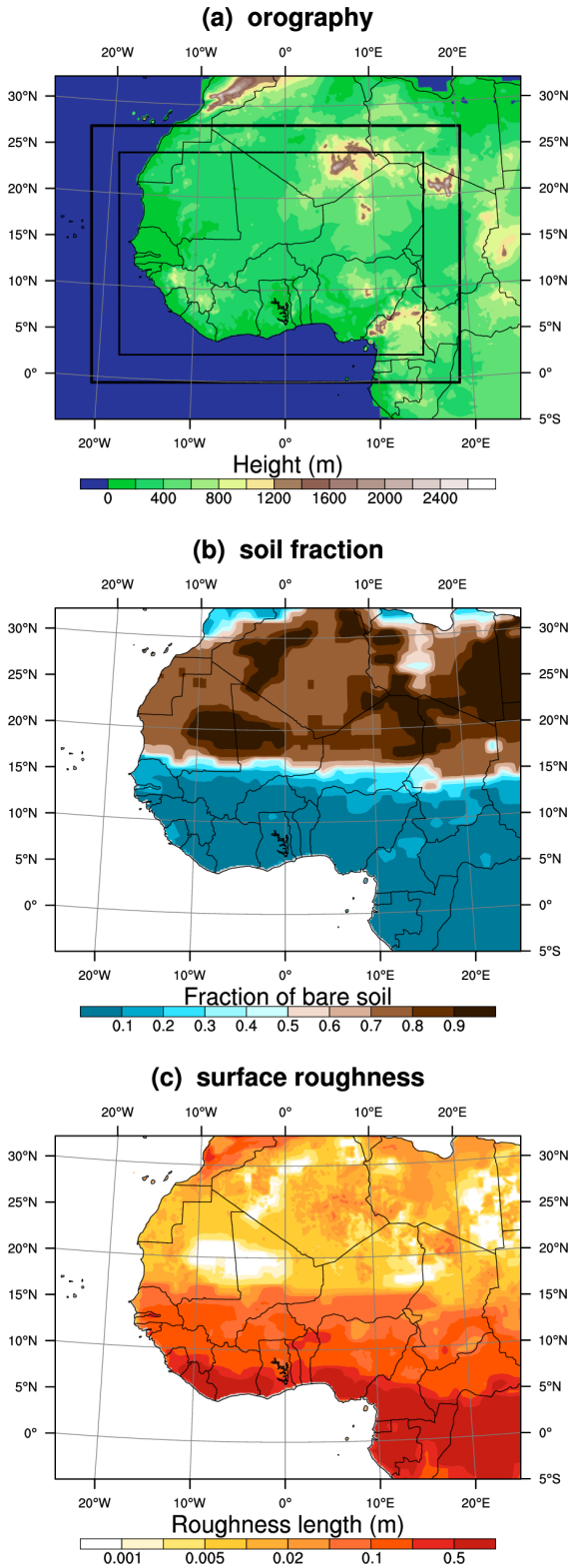
679 **Fig. 8.** Example of a cold pool outflow at 1500 UTC 1 July 2006 in the 4-km run: wind speed
680 (shading, in m s^{-1}) (a, c) and vertical velocity (shading, in m s^{-1}) (b, d), in horizontal (a, b)
681 and vertical cross-sections (c, d), showing the section-parallel wind (vectors above 3 m s^{-1}
682 according to the scale) and potential temperature (contours every 1 K) in the cross-sections.
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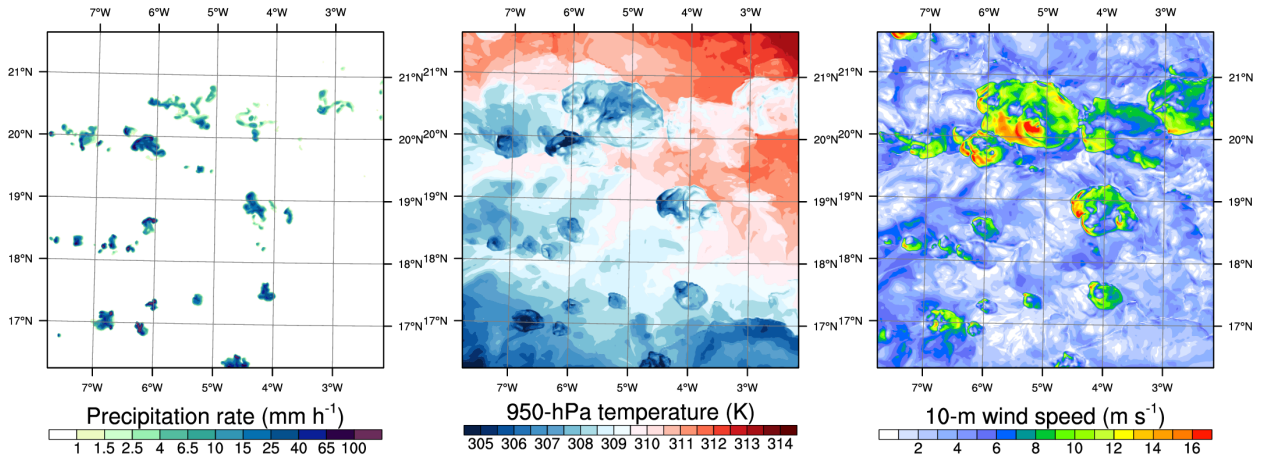
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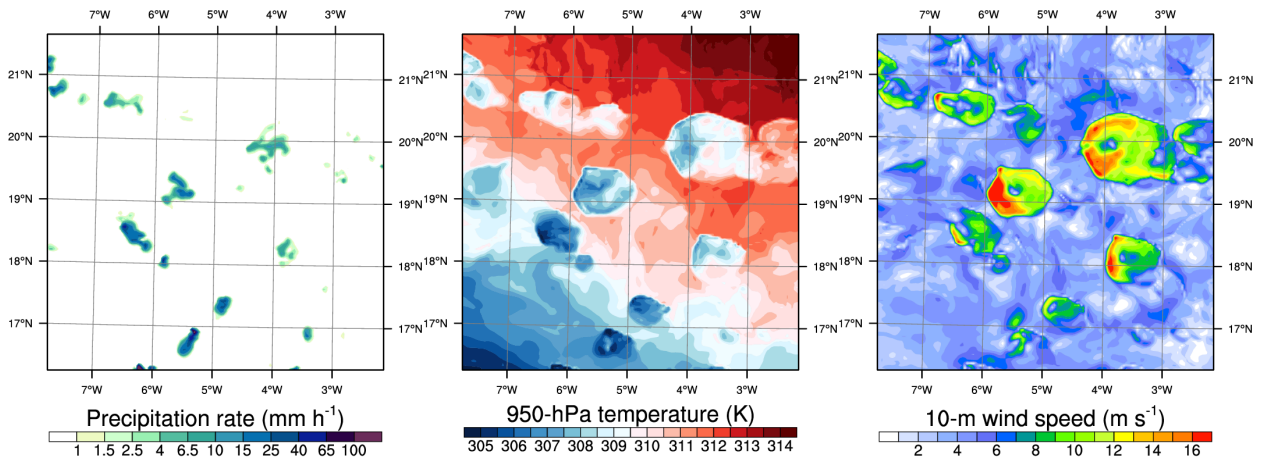


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 701 in (a) show the nested 4-km and 1.5-km domains, respectively.

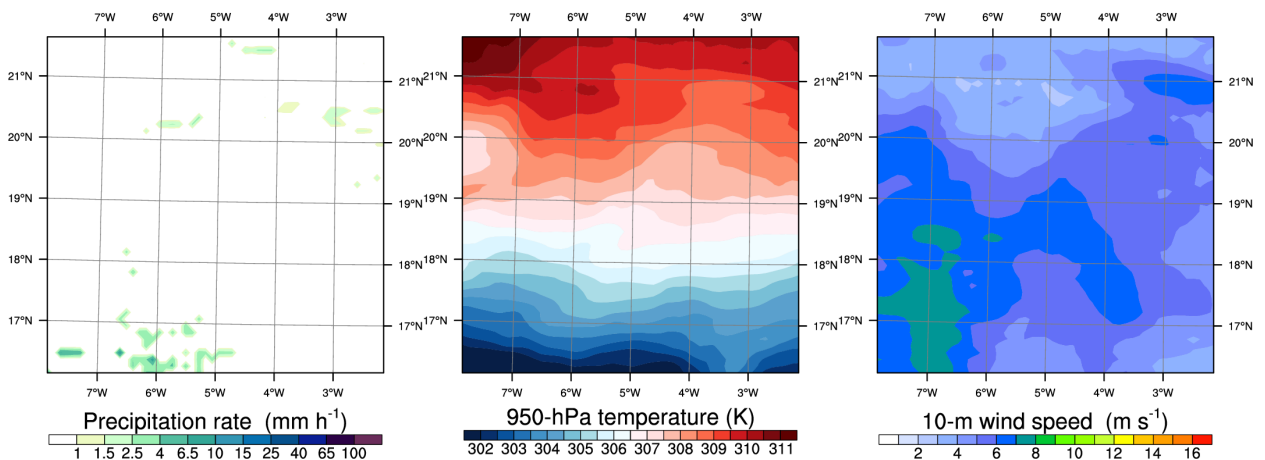
(a) 1.5 km: 1700 UTC 31 Jul 2006 (b) 1.5 km: 1700 UTC 31 Jul 2006 (c) 1.5 km: 1700 UTC 31 Jul 2006



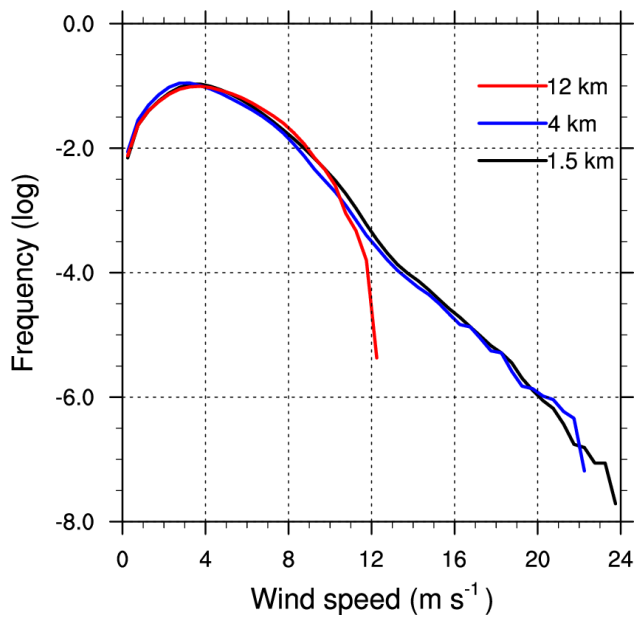
(d) 4 km: 1700 UTC 31 Jul 2006 (e) 4 km: 1700 UTC 31 Jul 2006 (f) 4 km: 1700 UTC 31 Jul 2006



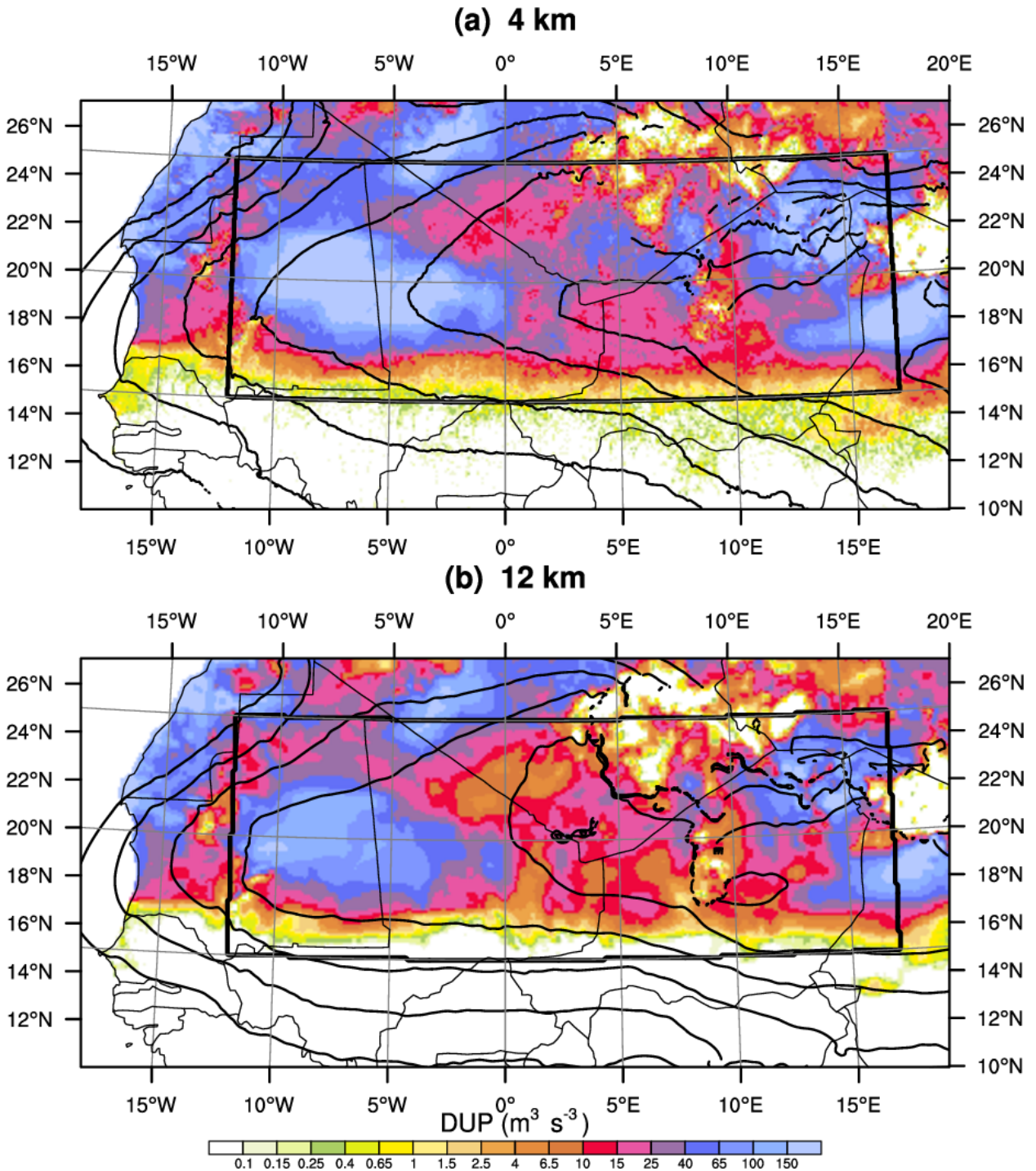
(g) 12 km: 1200 UTC 31 Jul 2006 (h) 12 km: 1200 UTC 31 Jul 2006 (i) 12 km: 1200 UTC 31 Jul 2006



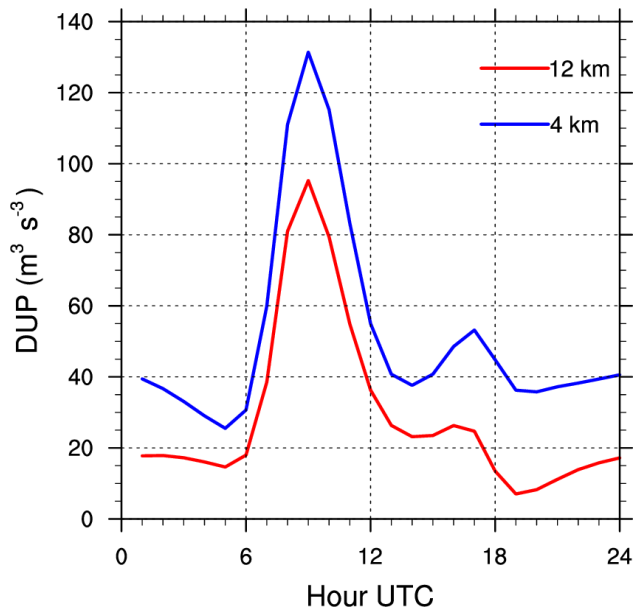
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 704 (in m s^{-1}) (c, f, i).



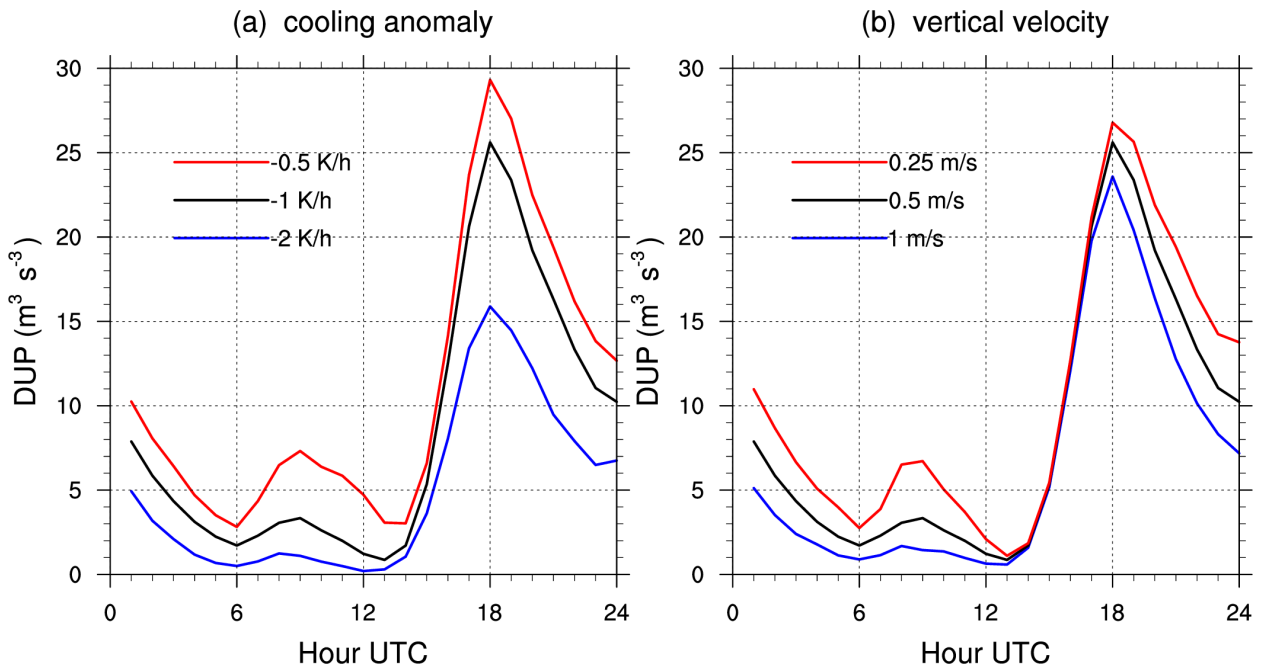
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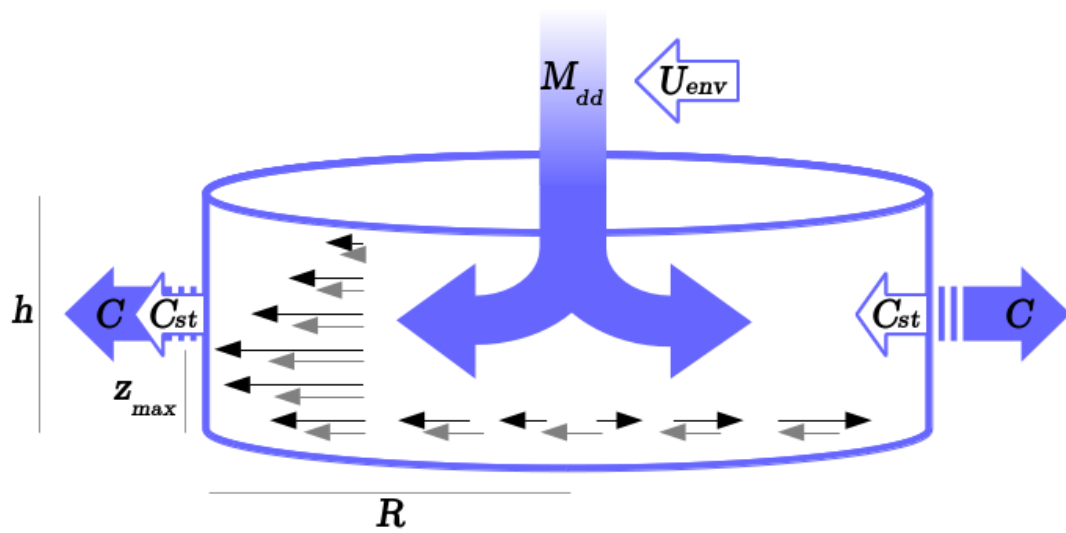
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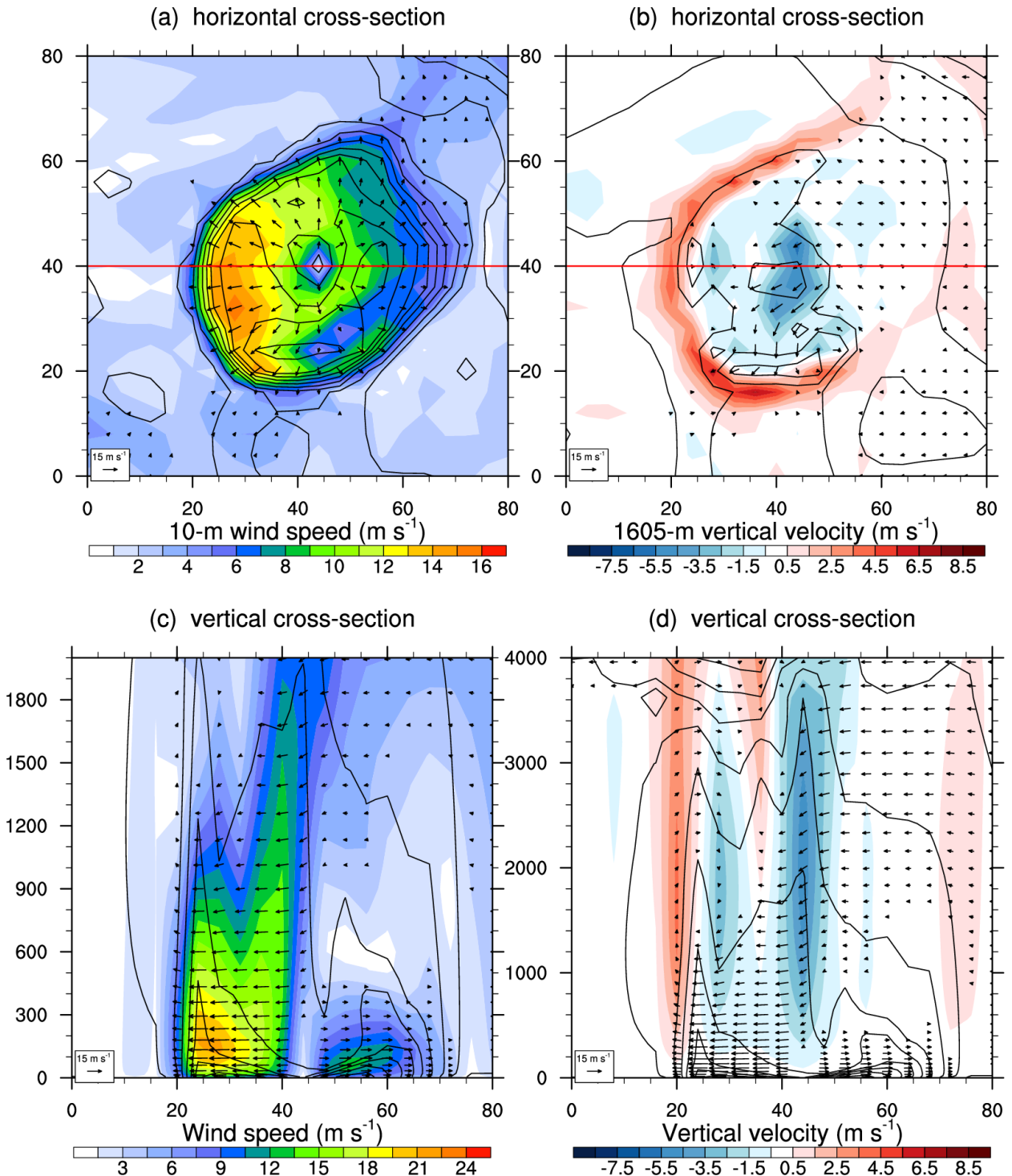
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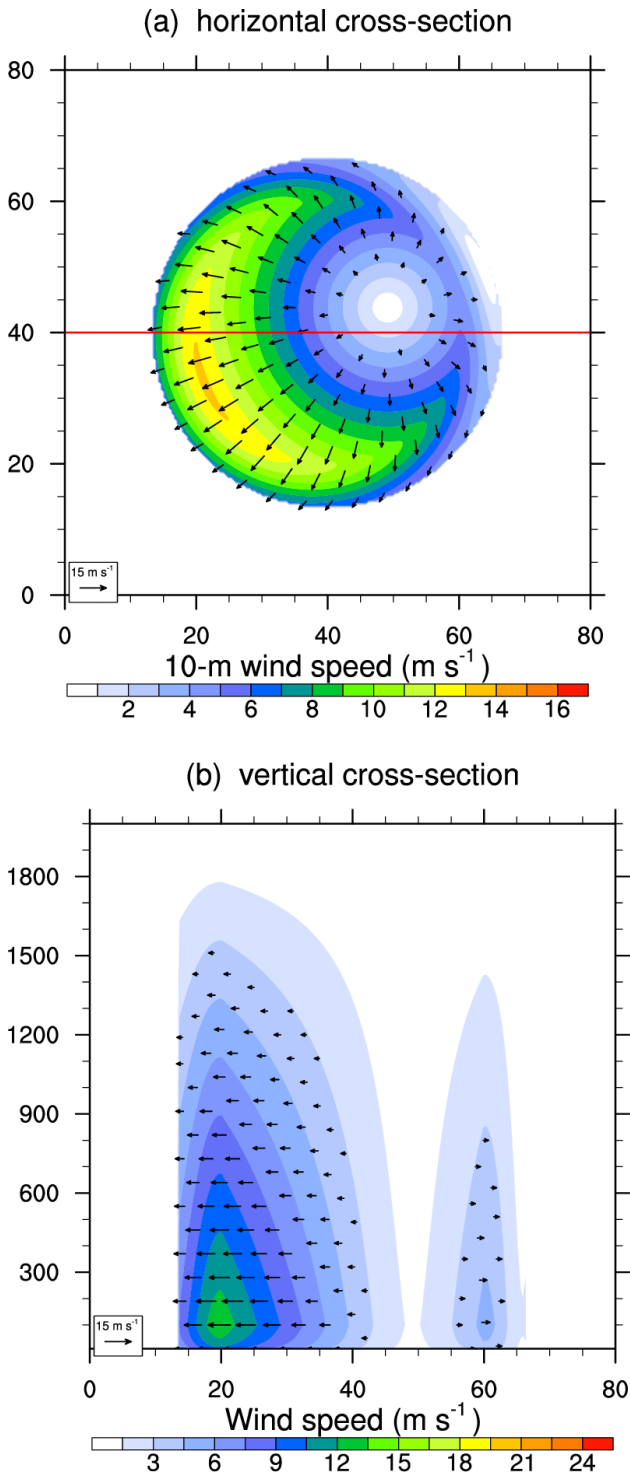
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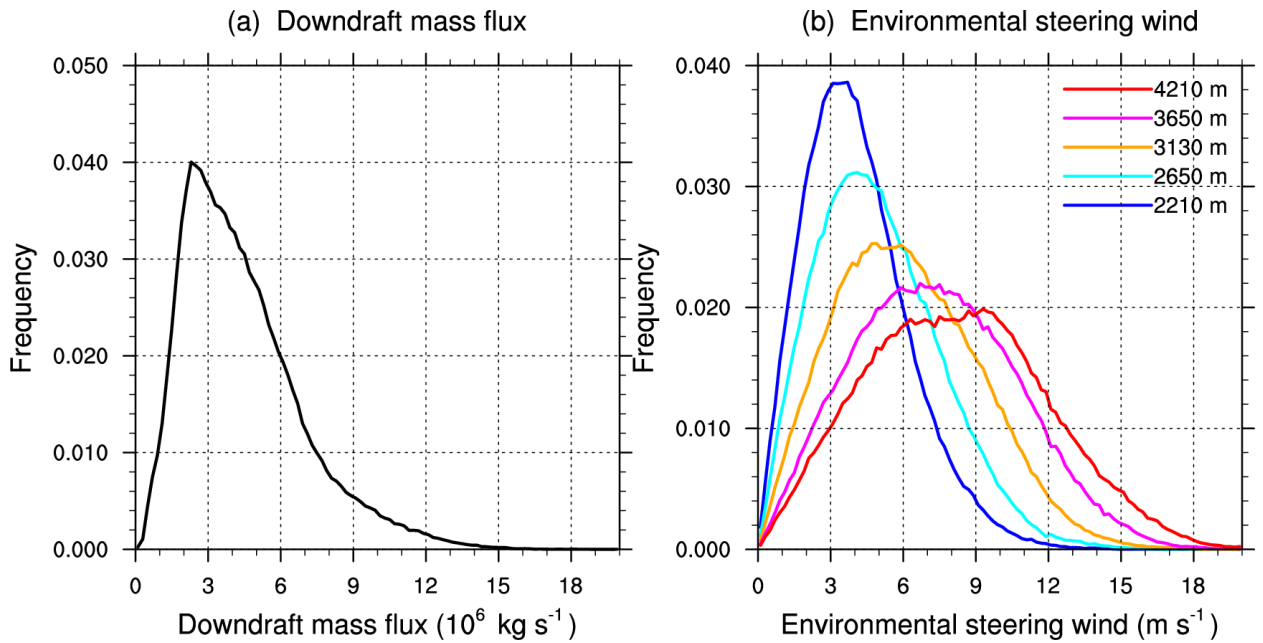
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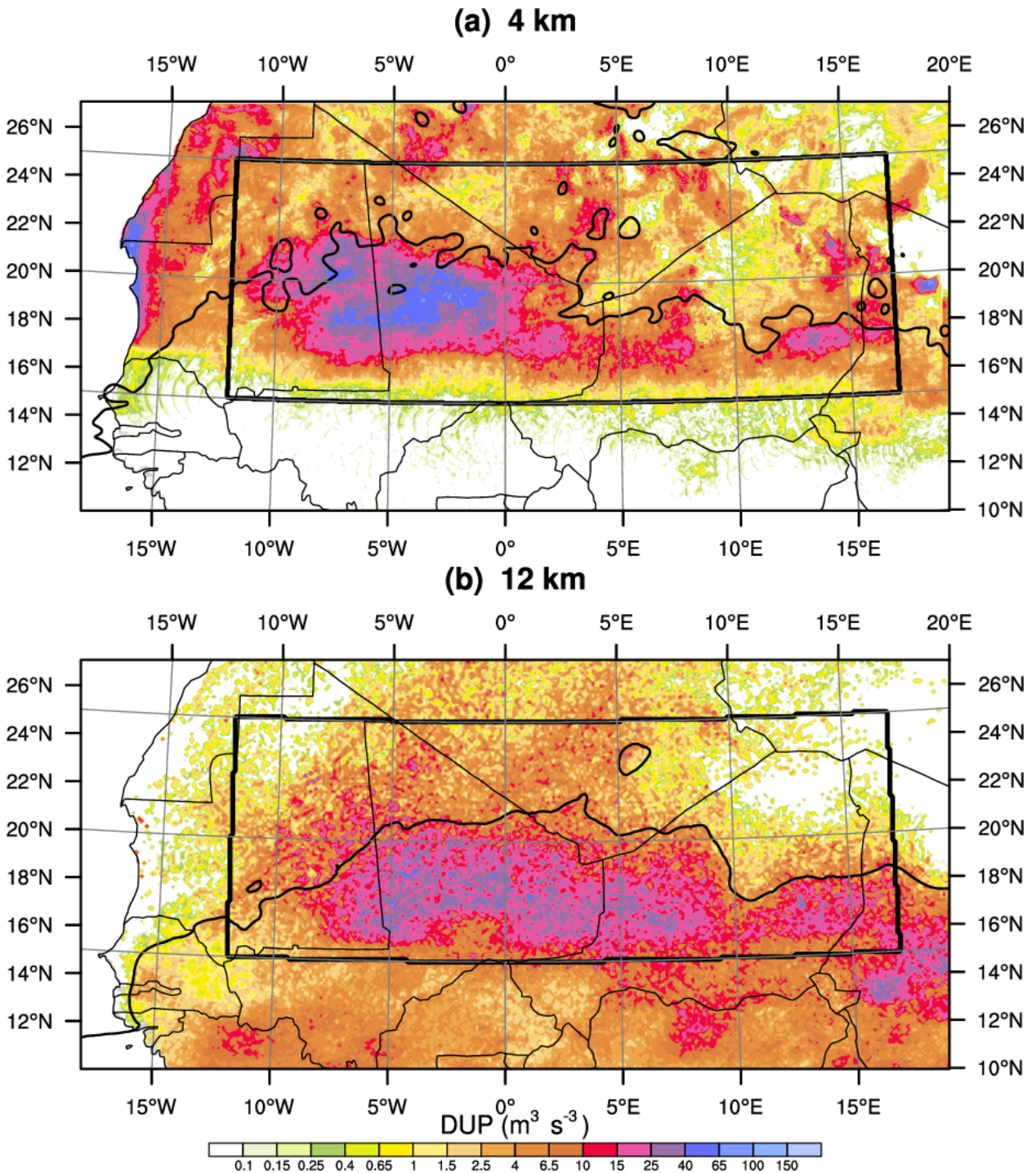
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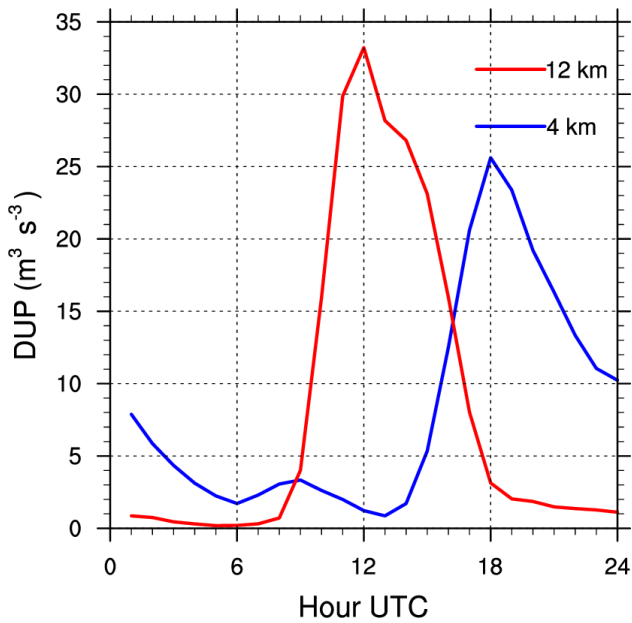
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