



Chemistry-climate model simulations of a mesospheric lower-thermospheric source of nitrous oxide

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SEE-Chem
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- Background
- N_2O production mechanisms
- First satellite observations
- Partial parameterisation
- Model setup
- Results
- Summary



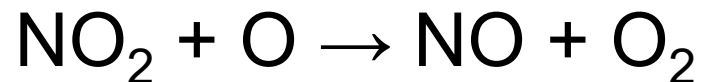
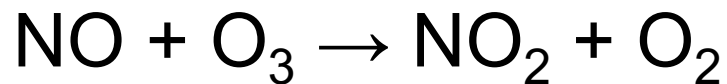
*Space Shuttle Endeavour
at the stratopause
Credit: NASA*

Ozone destruction

- N_2O is precursor of middle atmos. NO_y



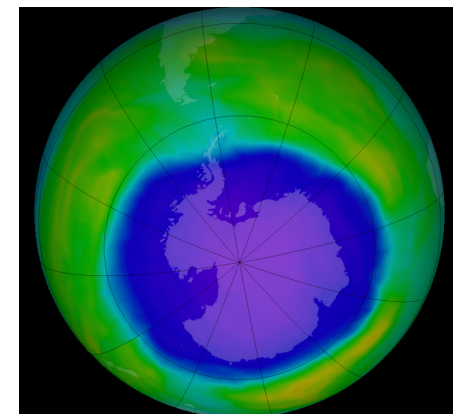
- NO_y destroys stratospheric O_3 :



~30% total ozone destruction

- Previously assumed only surface sources of N_2O
- **Mesospheric lower-thermospheric sources** now identified

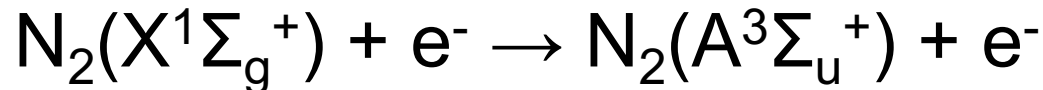
Ozone hole



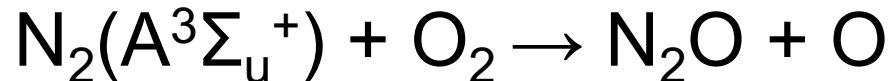
Credit: NASA

- Zipf and Prasad (1982) postulated:

- **Excited N₂** from **EEP** or **photoelectrons**



- Reaction with O₂



$\Delta E = 7.63 \text{ eV}$

ACE-FTS

- Produces N₂O at around **95km**
- But ~25 years without high-altitude infrared satellite observations to verify, until ACE-FTS



Credit: ACE – University of Waterloo

Production mechanisms (2/3)



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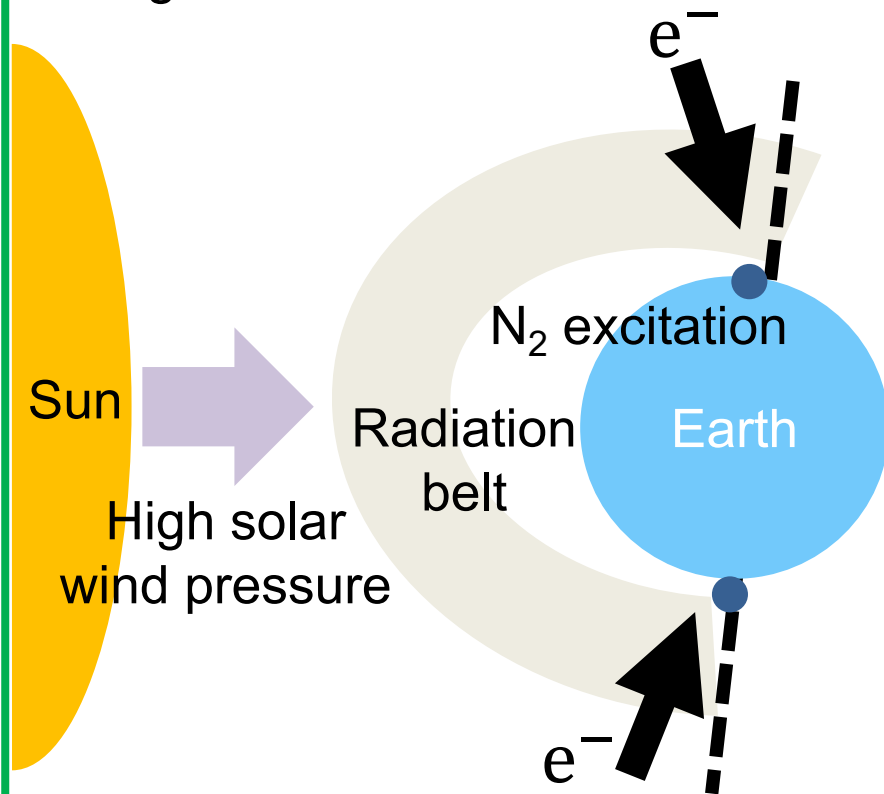
NERC
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Met Office

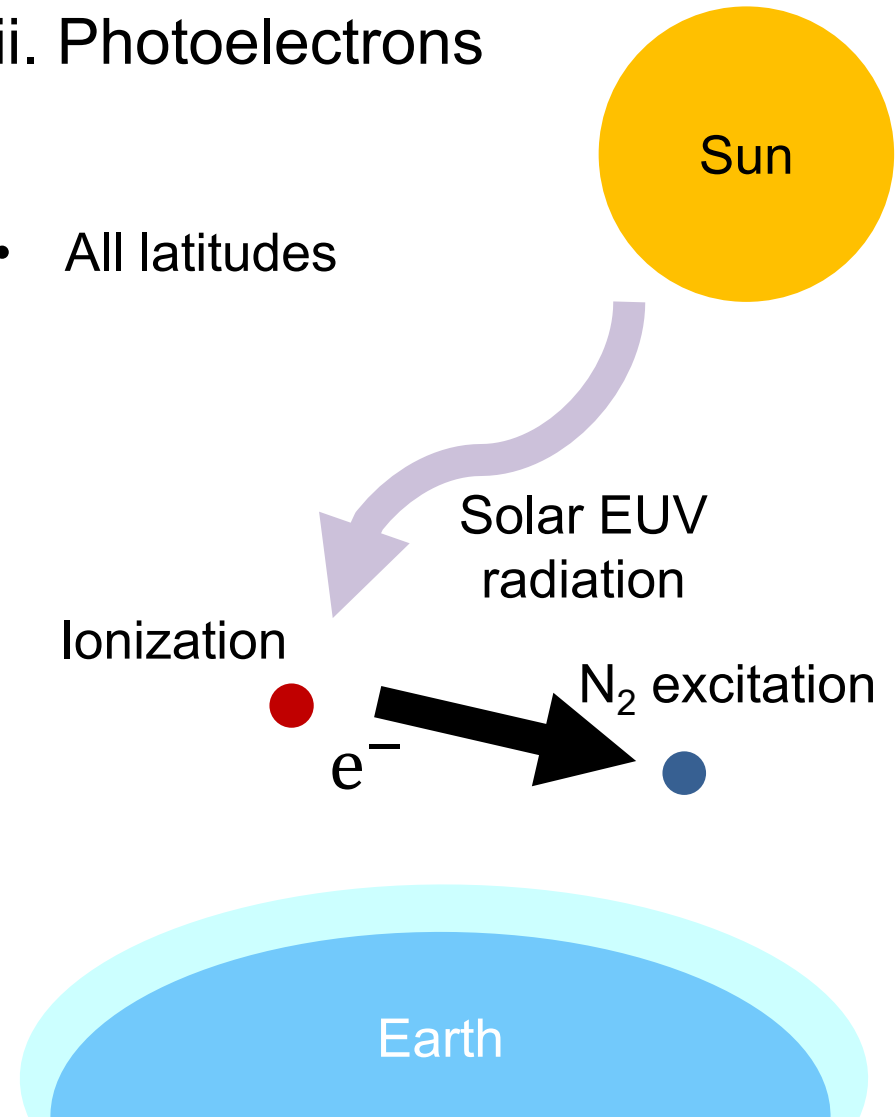
i. Energetic electron precipitation (EEP)

- High latitudes

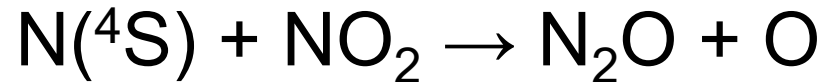


ii. Photoelectrons

- All latitudes



- Semeniuk et al. (2008) reported N_2O VMRs of ~5 ppbv at around 55 km using ACE-FTS data
 - Suggested production mechanism at around **75 km**:



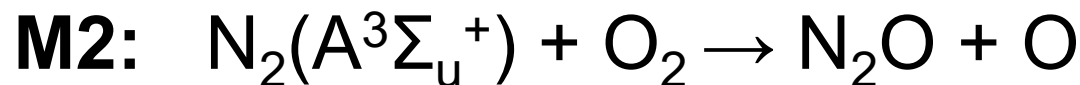
- Initially seen as upper atmospheric N_2O source

- *WACCM code already has...*



Semeniuk et al. (2008)

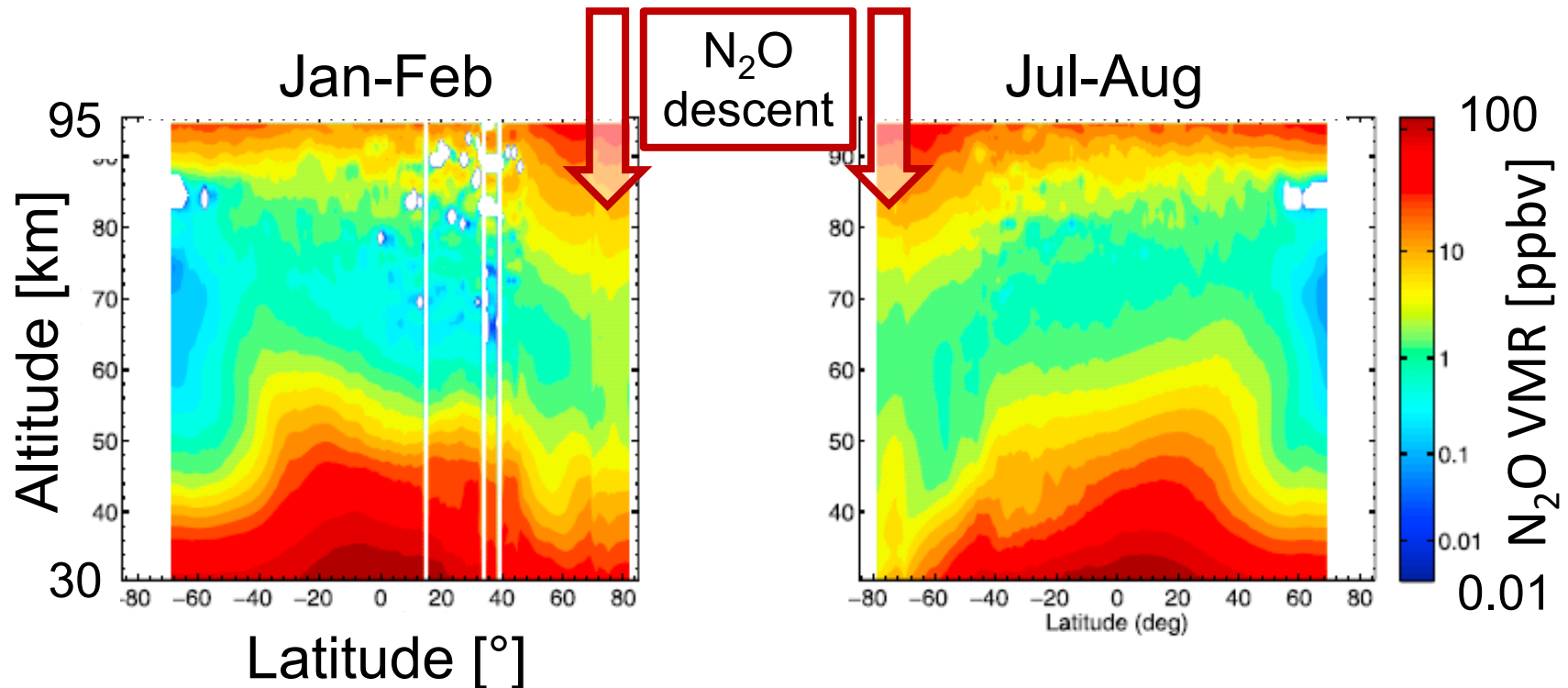
- *New mechanism...*



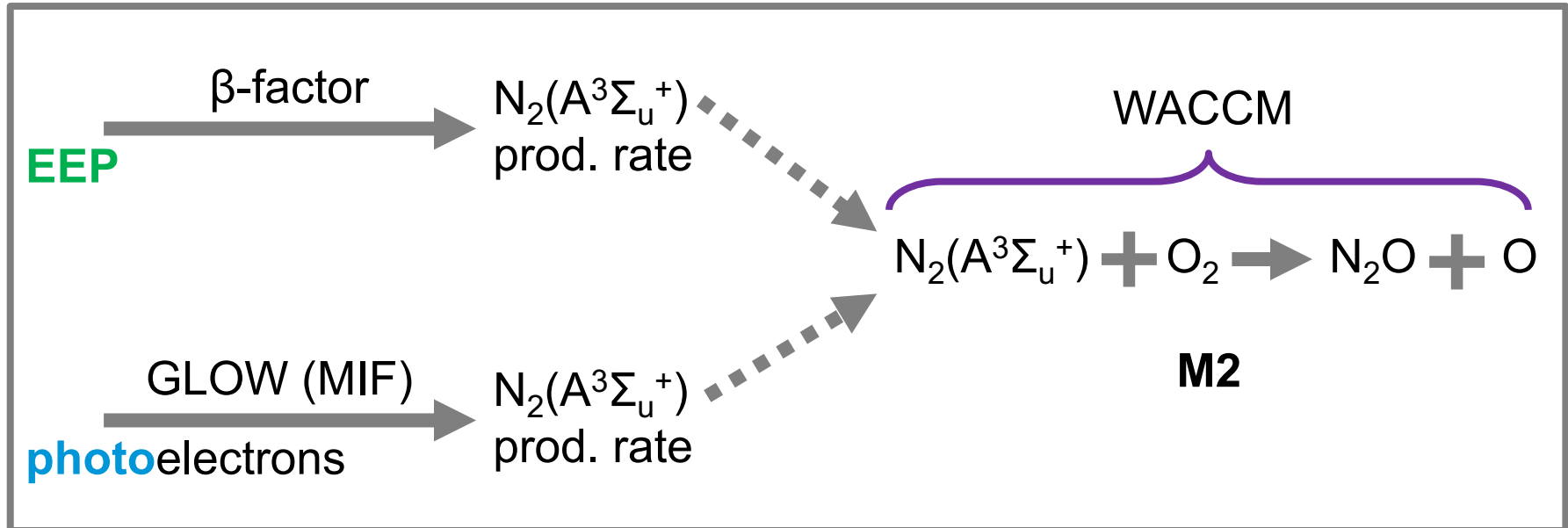
Zipf and Prasad (1982)

Date

- Sheese et al. (2016) provided first satellite obs. of 95 km N_2O with ACE-FTS
 - Reported maximum N_2O VMRs of 49 ppbv at 94.5 km



Inclusion of **M2** in WACCM:



EEP contribution

- Partially parameterise
- Efficiency (β) factor from steady-state assumption

Photoelectron contribution

- $N_2(A^3\Sigma_u^+)$ prod. rate from NCAR GLObal airglow model (*Solomon, 2017*)

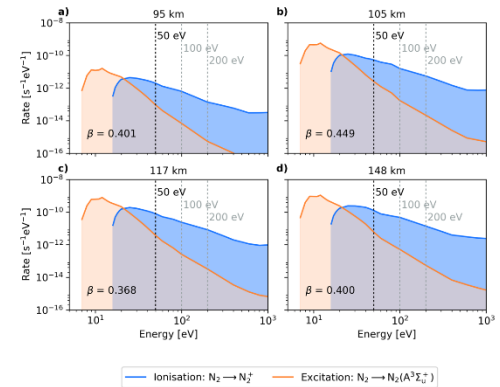
Steady state assumption:

$$\frac{d[\text{N}_2\text{O}]}{dt} = \text{Prod.} - \text{Loss} \cdot [\text{N}_2\text{O}]_{\text{obs}} = 0$$

$$\text{Total prod. rate: } \text{Prod.} = \left(\text{Prod.}_{\text{photo}} \cdot \alpha \right) + \left(\text{Prod.}_{\text{EEP}} \cdot \alpha \right)$$

Verified by second estimate based on first principles:

- *Lab. measurements of integrated cross section (ICS)*
- *Excitation of $\text{N}_2 \rightarrow \text{N}_2(\text{A}^3\Sigma_u^+)$ and the ionisation of N_2*
- *Convolved with auroral electron energy spectra*
- $\beta \approx 0.4$



$$\beta = \frac{\text{Loss} \cdot [\text{N}_2\text{O}]_{\text{obs}} - \left(\text{Prod.}_{\text{photo}} \cdot \alpha \right)}{\left(I_{\text{MEE}} + I_{\text{aur}} \right) \cdot \alpha} \approx 0.5$$

Model setup

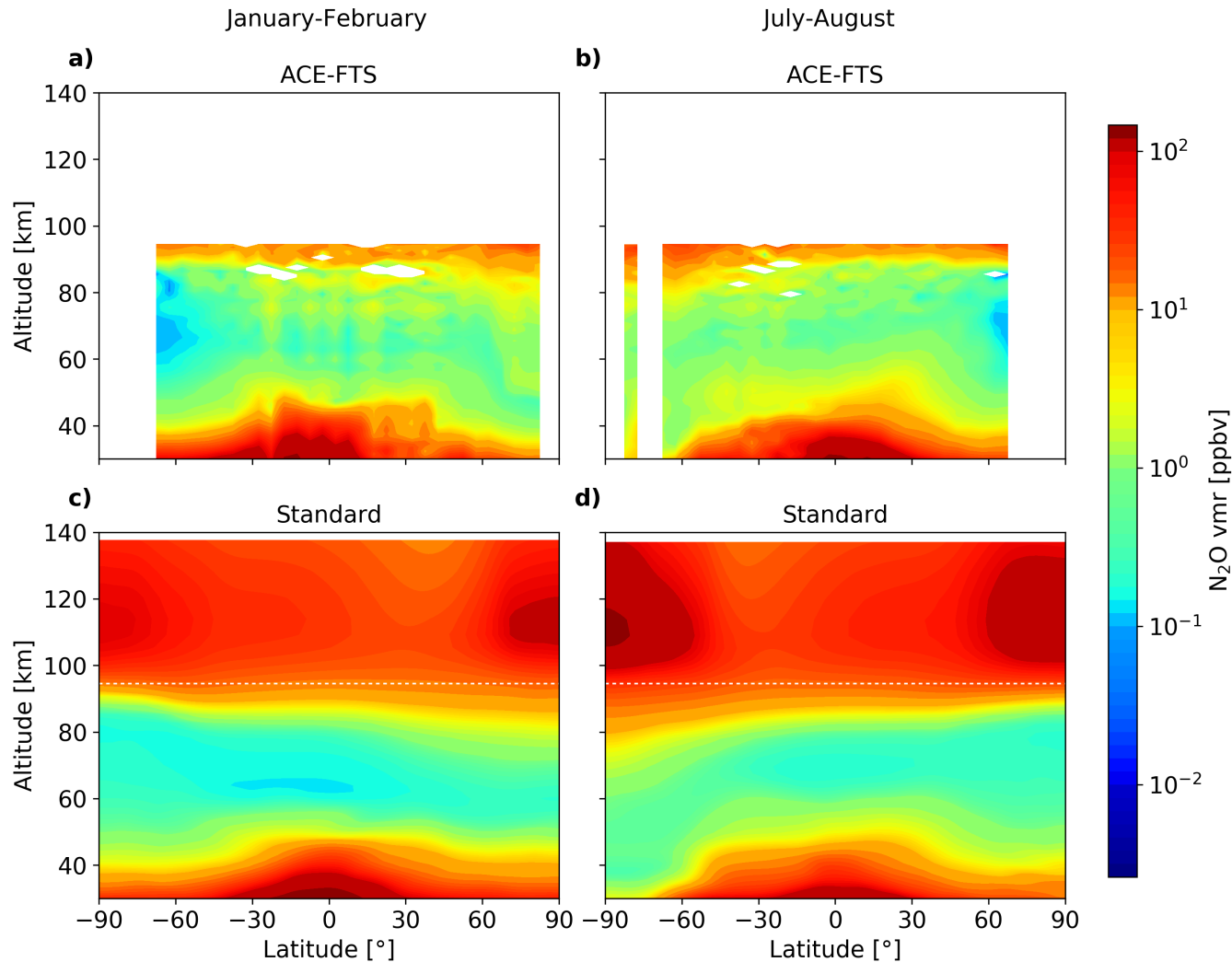
Simulation/ configuration	$N_2(A^3\Sigma_u^+)$ production via EEP	$N_2(A^3\Sigma_u^+)$ production via photoelectrons	(M2) $N_2(A^3\Sigma_u^+) + O_2$ $\rightarrow N_2O + O$	(M1) $N(^4S) + NO_2 \rightarrow$ $N_2O + O$	GLOW coupled
<i>Standard</i>	On	On	On	On	Yes
<i>Control_0</i>	Off	Off	Off	Off	No
<i>Control_1</i>	On	On	On	Off	Yes
<i>Control_2</i>	Off	Off	Off	On	No
<i>Sensitivity_E</i>	On	Off	On	On	No
<i>Sensitivity_P</i>	Off	On	On	On	Yes

WACCM changes:

- $N_2(A^3\Sigma_u^+)$ tracer added
- M2 and other depletion routes included

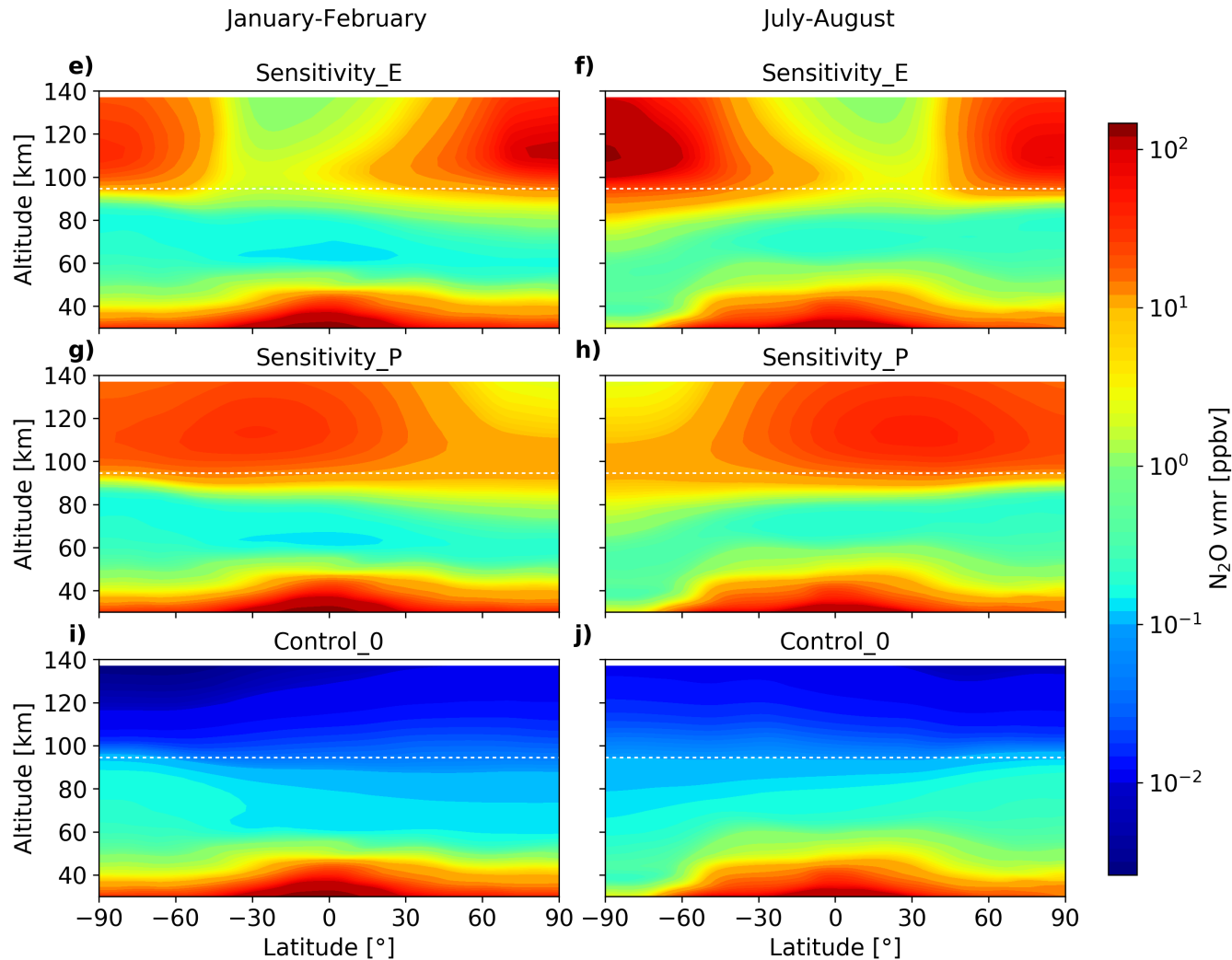
All simulations
completed over 2013

Latitude cross-sections (1/2)



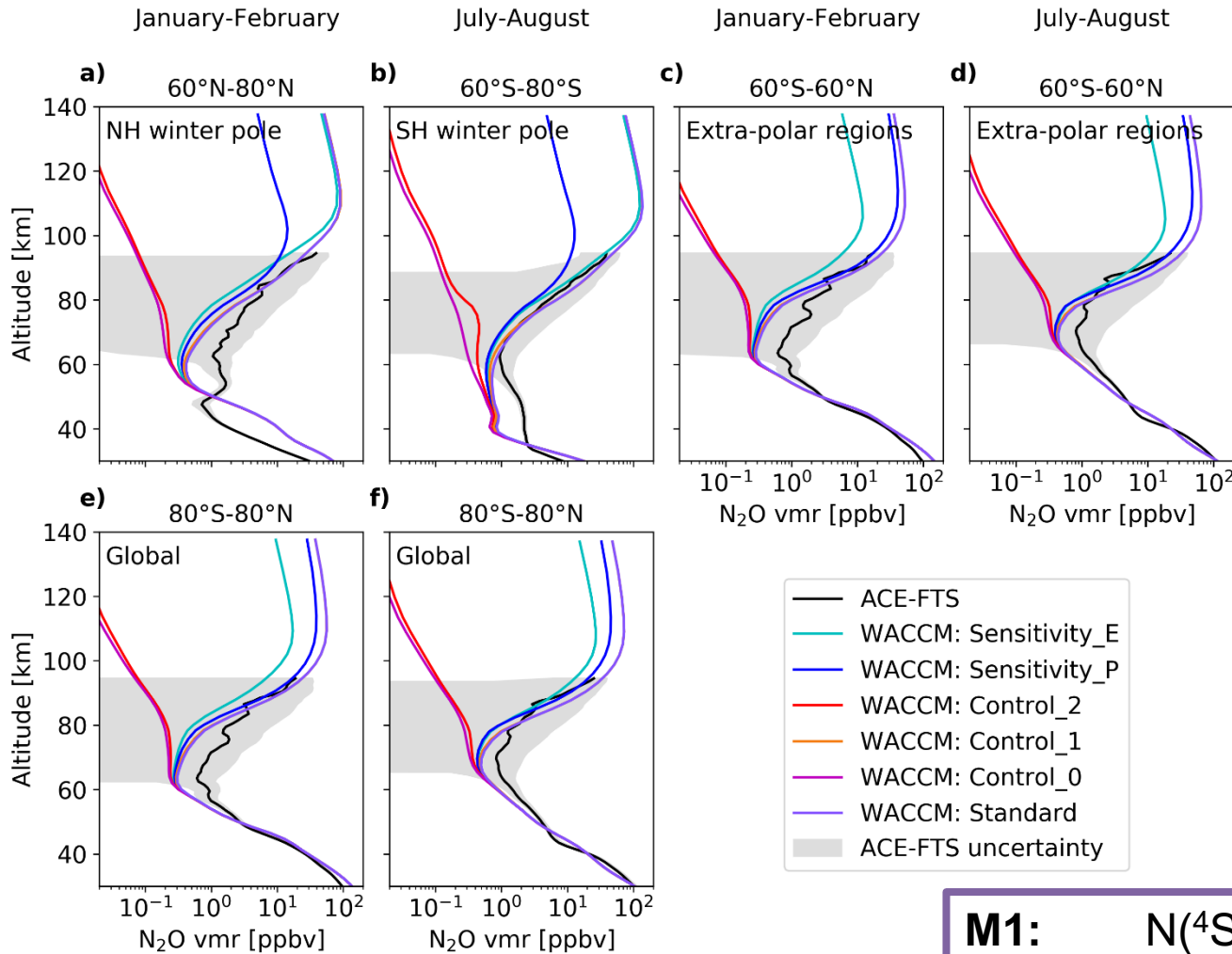
- 2D latitude cross-sections of N_2O ppbv
- Satellite (top), model (bottom)
- Note mixing ratios – not concentrations

Latitude cross-sections (2/2)

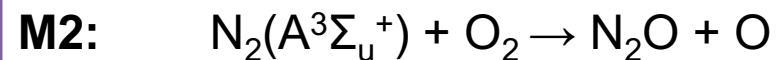
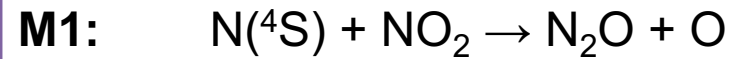


- **EEP** dominant at poles
- **Photoelectrons** dominant at mid and low latitudes
- Without *in-situ* N_2O prod. mixing ratios decrease exponentially with altitude

Altitude profiles



- Model matches satellite in MLT
- >99% N₂O prod. from **M2** above 80 km
- Max **M1** impact over SH winter pole (~20%)



- *In-situ* N₂O production in the mesosphere lower-thermosphere
- Excited N₂ from **EEP** or **photoelectrons** reacts with O₂ to form N₂O
- Added into WACCM via partial parametrisation
- Model replicates observed vertical, latitudinal and seasonal profile
- **EEP** dominant at poles, **photoelectrons** dominant at mid and low latitudes

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