# Towards imaging flow at the base of the mantle with seismic, mineral physics and geodynamic constraints

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#### 6 Abstract

Perhaps the least ambiguous signal that the mantle is convecting comes from observa-7 tions of seismic anisotropy—the variation of wave speed with direction—which must arise 8 due to the ordering of material as deformation occurs. Therefore significant effort has 9 been made over many years to infer the direction and nature of mantle flow from these 10 data. Observations have focussed on the boundary layers of the mantle, where deforma-11 tion is expected to be strongest and where anisotropy is usually present. While prospects 12 for mapping flow seem good, the lack of knowledge of several key issues currently holds 13 progress back. These include the cause of anisotropy in the lowermost mantle, the causative 14 material's response to shear, and the single-crystal or -phase seismic properties of the 15 causative materials. In this chapter we review recent observations of lowermost mantle 16 anisotropy, constraints on mineral elasticity and deformation mechanisms, and challenges 17 in linking geodynamic modelling with seismic observations. 18

#### <sup>19</sup> 1 Introduction

Seismic anisotropy, i.e. the variation of seismic velocity with propagation direction 20 and polarisation, is observed in a number of regions within the Earth. The strength of 21 anisotropy is particularly strong in the crust, at the top and the bottom of the mantle, 22 and in the inner core (Mainprice, 2007). In the upper mantle, observed seismic anisotropy 23 has been used to map asthenospheric flows (Becker and Lebedev, this volume) and un-24 derstand slab dynamics (Huang and Zhao, this volume). In the lowermost mantle, un-25 derstanding anisotropy in terms of flow is more elusive, as seismic observations are sparse, 26 and mineral physical constraints more uncertain. The dynamics of the lowermost man-27 tle are of particular interest, as they reflect the lower thermal and mechanical bound-28 ary layer of the convecting mantle. Mapping flow directions in this region would signif-29 icantly help our understanding of the role of this boundary in mantle dynamics, and more 30 specifically the role of the large low-shear velocity provinces ('LLSVPs') (Rudolph et al., 31 this volume). 32

Albeit challenging, significant efforts have been made to use seismic anisotropy to understand the underlying crystal preferred orientations (CPO) and flow directions (e.g., Karato, 1998). This is based on the assumption that flow, and the internal crystallographic deformation mechanisms that accommodate the flow, cause alignment of intrinsically anisotropic crystals. This chapter offers a condensed review of seismic observations and mineral phys-

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ical and geodynamical constraints on seismic anisotropy, and for a more in-depth review 38 we refer to Nowacki et al. (2011) and Romanowicz & Wenk (2017). Here, we focus on 39 the endeavours, mainly over the last decade, to the these disciplines together and map 40 flow directions in the lowermost mantle, and the specific challenges posed when compar-41 ing these results to seismic observations.

#### 2 Observational constraints on lowermost mantle flow 43

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## 2.1 Global tomographic models

A number of tomographic modellers invert for seismic anisotropy in the lowermost 45 mantle. Inverting the full anisotropic elastic tensor (i.e., all 21 parameters) is unfeasi-46 ble. The only component of anisotropy generally inverted for in the lowermost mantle 47 is the velocity difference between horizontally polarised shear velocity,  $V_{SH}$ , and verti-48 cally polarised shear velocity,  $V_{SV}$ . This component of anisotropy is named radial anisotropy 49 (or vertically transverse isotropy) and the degree of anisotropy can be expressed by the 50 value  $\xi = V_{SH}^2/V_{SV}^2 = C_{66}/C_{44}$ , where C is the Voigt matrix representation of elas-51 ticity and the 3-axis is vertical. 52

For the upper mantle, radial anisotropy is well constrained due to the unique sen-53 sitivities of the two types of surfaces waves (Becker and Lebedev, this volume). For the 54 lower mantle, one or both of normal modes and body waves must be used. Normal mode 55 inversions for 1D radial models show no significant component of  $\xi$  (Beghein et al., 2006; 56 de Wit & Trampert, 2015). 3D tomographic models show a mainly isotropic lower man-57 tle with lateral variations in  $\xi$  on the order of 0.97–1.03 in the lowermost mantle (e.g., 58 Moulik & Ekstrom, 2014; Auer et al., 2014; Chang et al., 2015; French & Romanowicz, 59 2015). There are strong differences between these models, some of which can be attributed 60 to their treatment of the crust in the inversion, which is shown to affect the radial anisotropic 61 signature of the lowermost mantle (Ferreira et al., 2010; Panning et al., 2010). In gen-62 eral, a geographical trend emerges where  $\xi > 1$  (equivalently,  $V_{SH} > V_{SV}$ ) in regions 63 with fast shear wave velocity interpreted to be slab graveyards, and opposite signature 64 of  $\xi < 1$  is seen in regions of slow shear wave velocity, i.e. the LLSVPs. In Figure 1 this 65 relationship is illustrated with histograms of  $\xi$  values for the fast and slow regions as in-66 terpreted by Cottaar & Lekić (2016), as well as a vote map of  $\xi$  values, which at each 67 point at 2800 km depth shows the count of all tomography models which have a value 68

of  $\xi$  above or below 1. While all models show a significant shift in their histograms for 69 the fast and slow region, the shifts between their mean values is small, with the largest 70 shift of 1.1% in SEMUCBwm1, and the smallest shift of 0.38% in SAVANI. The vote map 71 also suggests a relationship between  $dV_S$  and  $\xi$ . Interpreting this general trend should 72 be done with caution as any relationship between  $dV_S$  and  $\xi$  could be an artefact of the 73 inversion, specifically the negative  $\xi$  values appear prone to be leakage of the slow isotropic 74 velocities (Chang et al., 2015). One thing that is interesting to note from the vote map 75 is the smaller slow shear region beneath the Ural mountains, dubbed the Perm anomaly 76 (Lekić et al., 2012), does not correlate directly with a signal of  $\xi < 1$ , but a small anomaly 77 showing  $\xi < 1$  appears offset to the south. 78

Studies are even more limited in constraining P wave radial anisotropy ( $\phi = V_{PV}^2/V_{PH}^2 =$ 79  $C_{33}/C_{11}$ ). Global studies constraining 1D radial P wave anisotropy show no agreement 80 in the likely signature (Beghein et al., 2006; de Wit & Trampert, 2015). Global 3D stud-81 ies have often applied an assumed scaling between the S and P wave anisotropy as a start-82 ing model. When they do include P wave radial anisotropy independently (Soldati et al., 83 2003; Tesoniero et al., 2016), they judge their results not to be robust. Inversions using 84 body waves are heavily under-constrained (Boschi & Dziewoński, 2000). The synthetic 85 study of P. J. Koelemeijer et al. (2012) shows general sensitivity of normal modes to P 86 wave anisotropy, although it also predicts it is sensitive to trade-offs. 87

A potential way forward for global studies lies in the understanding of a third anisotropic parameter,  $\eta = C_{13}/(C_{11} - 2C_{44})$ , which is related to the S and P velocities at intermediate incidence angles. de Wit & Trampert (2015) show that this parameter has a robust 1D signature of  $\eta < 1$  across the lower 1000 km of the mantle. Kawakatsu (2015) suggests a rewrite of  $\eta$  for a more physical relationship with wave incidence angle, and shows that constraining this parameter,  $\eta_{\kappa}$ , can help resolve whether horizontally isotropic layers could cause the observed anisotropy.

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### 2.2 Regional body wave observations

Locally, deep mantle seismic anisotropy can be observed through shear wave splitting of body waves. One of the main challenges is to determine the relative contribution to splitting from the uppermost and lowermost mantle, and whilst often assuming the rest of the lower mantle is isotropic. A general approach is to use two seismic phases with

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comparable ray paths across the upper mantle, while one reference phase has a different or no ray path across the lowermost mantle. Ideally the reference phase shows no or minimal splitting and all splitting in the other phase can be attributed to the lowermost mantle. Otherwise corrections for splitting from the upper mantle need to be applied to attribute splitting to the lower mantle (e.g., Wookey et al., 2005).

One potential set consists of the ScS and S phases (Figure 2). For the distance range 114 of  $60-85^{\circ}$ , the S phase turns above the lowermost mantle, while the ScS phase samples 115 the lowermost mantle (e.g., Lay & Young, 1991; Wookey & Kendall, 2008). A second set 116 is the SKS and SKKS phases (at epicentral distances  $108-122^{\circ}$ ), for which the ray paths 117 exit the core at different locations and converge across the mantle (e.g., Niu & Perez, 118 2004; Wang & Wen, 2007; M. Long, 2009). SKS–SKKS pairs have the additional advan-119 tages over S–ScS that in an isotropic or radially-anisotropic mantle they exit the core 120 purely polarised along the SV component, and any anisotropy along the down-going leg 121 of the path can be ignored. Their disadvantage is that both phases can accrue splitting 122 in the lowermost mantle and in the upper mantle. At times it is difficult to retrieve split-123 ting parameters and these phases are usually only used to highlight discrepant phase pairs 124 (e.g., Deng et al., 2017). 125

Lastly,  $S_{diff}$  phases (at 100–130° distance) are compared either to S/ScS at shorter distances, or to SKKS (or SKS) at longer distances (e.g., Kendall & Silver, 1996; Vinnik et al., 1998). The SV component of the diffracted wave attenuates much faster than the SH along the core-mantle boundary, which means  $S_{diff}$  at large distances (>~ 120°) becomes a purely SH polarised wave, and any splitting can be attributed to the upgoing leg of the ray path (Cottaar & Romanowicz, 2013).

In all cases, caution is required in interpreting body-wave observations if modelled 132 using approximate methods such as ray theory, since shear waves at the base of the man-133 the have a large region of finite-frequency sensitivity. For  $S_{diff}$ , travel time differences be-134 tween SH<sub>diff</sub> and SV<sub>diff</sub> can arise for purely isotropic models, especially with strong isotropic 135 velocity gradients as one might expect due to the thermal boundary layer, due to dif-136 ferent finite-frequency sensitivity of the two components along the boundary (Maupin, 137 1994; Komatitsch et al., 2010; Borgeaud et al., 2016; Parisi et al., 2018). ScS suffers to 138 a lesser extent from finite-frequency effects in 1D models, but ray theoretical interpre-139

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tations can badly misrepresent the strength and orientation of anisotropy when lateral
variations in anisotropy may exist (Nowacki & Wookey, 2016).

The different phases have different sensitivity to the anisotropic tensor due to their 142 propagation angle and the length of their propagating path across the lowermost man-143 tle (Figure 2). S<sub>diff</sub> has long horizontal propagation paths in the mantle, and therefore 144 good sensitivity to radial anisotropy. SKS propagates at sub-vertical angles  $(18^\circ - 33^\circ)$ 145 across the lowermost mantle, so splitting is caused by the component of azimuthal anisotropy, 146 i.e. variation of wave speed in the horizontal plane. SKKS  $(40^\circ-50^\circ)$  and ScS  $(62^\circ-78^\circ)$ 147 propagate at intermediate angles, and are sensitive to tilted anisotropy. While early stud-148 ies focused mainly on constraining the radial anisotropic component (e.g., Young & Lay, 149 1990; Matzel et al., 1996; Garnero & Lay, 1997), recent studies interpret their observa-150 tions as tilted anisotropy, the main component constrained when accounting for the in-151 cidence angles in the lowermost mantle (e.g., Thomas et al., 2007; Wookey & Kendall, 152 2008; Nowacki et al., 2010). 153

One additional, unique type of observation worth mentioning are polarity observations of phases bouncing off of the so-called D" discontinuities in the lowermost mantle (Thomas et al., 2011; Cobden & Thomas, 2013; Creasy et al., 2019; Pisconti et al., 2019). Azimuthal variations in the polarity measurements suggest these are sensitive to underlying anisotropy. As observations can be applied to S and P reflections ('SdS', and 'PdP'), they are to our knowledge the only body wave studies that have resolved a component of both S and P wave anisotropy for a single location.

Most observational studies focus on a single observational method, as well as a single azimuthal direction. To sufficiently constrain anisotropy in a single location to uniquely interpret flow direction, multiple techniques need to be combined (Creasy et al., 2019). Efforts have been made to target a single region from multiple angles using ScS (Nowacki et al., 2010; Wookey & Kendall, 2008) and polarisation measurements (Thomas et al., 2011), as well as combining multiple angles with multiple body wave phases (Ford & Long, 2015; Creasy et al., 2019; Wolf et al., 2019).

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### 2.3 Observed regional anisotropy

This is not an exhaustive overview of body wave studies and for a full table of studies we refer to Romanowicz & Wenk (2017). Here we highlight consistencies across these

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studies, mainly focusing on more recent studies which benefit from increased coverage 171 by seismic arrays. Regional body wave studies largely agree with tomographic models 172 on geographical trends in radial anisotropy, i.e.  $\xi > 1$  where isotropic velocities are fast, 173 and  $\xi < 1$  where isotropic velocities are slow (e.g., Wookey & Kendall, 2007; Kawai & 174 Geller, 2010). Models interpreting tilted anisotropy have overwhelmingly sampled isotrop-175 ically fast areas and many find a sub-horizontal fast axis and thus a component of  $\xi >$ 176 1 (e.g., Thomas et al., 2007; Garnero et al., 2004; Wookey & Dobson, 2008; Nowacki et 177 al., 2010), while several studies find a fast axis which is tilted from the horizontal by around 178 45° (Wookey et al., 2005; Cottaar & Romanowicz, 2013), which is not compatible with 179 radial anisotropy. Particularly, regions just outside of LLSVPs appear to have strong and 180 variable anisotropy, as is observed along the boundaries of the African LLSVP (Wang 181 & Wen, 2007; Cottaar & Romanowicz, 2013; Lynner & Long, 2014; Grund & Ritter, 2019; 182 Romanowicz & Wenk, 2017; Reiss et al., 2019), the Pacific LLSVP (Deng et al., 2017), 183 and the Perm Anomaly (M. D. Long & Lynner, 2015). These observations show stronger 184 anisotropy outside of the LLSVP and little to no anisotropy within the LLSVP, both in 185 terms of tilted anisotropy (Cottaar & Romanowicz, 2013) and in terms of azimuthal anisotropy 186 (e.g., Lynner & Long, 2014; Grund & Ritter, 2019). A change in sign from  $\xi > 1$  to  $\xi <$ 187 1 is also observed towards the base of the Icelandic plume (Wolf et al., 2019). 188

While some consistency emerges on the types of anisotropy, and correlations with 189 isotropic velocities, uncertainty lies in the strength of anisotropy observed. Tomographic 190 models contain radial anisotropy on the order of several %, and amplitudes vary between 191 models (see Figure 1). Local observations interpret tilted anisotropy of 0.8-1.5% across 192 a layer of 250 km beneath North America (Nowacki et al., 2010) and up to 8% across 193 150 km beneath the Antarctic Ocean (Cottaar & Romanowicz, 2013). Such variations 194 could represent true geographical observations, but biases could also occur as propaga-195 tion angles used might not be optimal to observe the strongest splitting or assumed layer 196 thicknesses. In these two example studies, interpreted amplitudes might also differ as 197 one is interpreted ray-theoretically (Nowacki et al., 2010) and one by forward modelling 198 (Cottaar & Romanowicz, 2013). Potentially, consistently constrained relative amplitudes 199 in splitting might help map lateral variations in flow strength or direction. 200

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### <sup>210</sup> **3** Forward modelling

To provide synthetic tests for the hypothesis that anisotropy is caused by crystal 211 preferred orientation (CPO), multi-disciplinary models are built that span many spa-212 tial scales (see flow chart in Figure 3). Geodynamic models provide maps of strain across 213 10s to 1000s of km. The strain observed is accommodated on the micro scale by defor-214 mation mechanisms in a set of crystals, assuming a degree is accommodated by dislo-215 cation glide to create preferential orientation. The individual elastic constants of each 216 of the deformed set of crystals are averaged using their orientations, giving the fully anisotropic 217 tensor for a single location. This process needs to be repeated for many locations, to pro-218 vide an anisotropic model with signatures that can be observed over 10s or 100s of km 219 by seismic waves. Here we explain the main choices and assumptions made in these mod-220 221 els.

While we focus on the hypothesis that CPO is the cause of seismic anisotropy in 222 the lowermost mantle, studies have forward modelled the potential of shape preferred 223 orientation (SPO) as well. SPO anisotropy is caused by layering or inclusions of strongly 224 heterogeneous (but potentially intrinsically isotropic) material (Kendall & Silver, 1998; 225 Hall et al., 2004; Creasy et al., 2019; Reiss et al., 2019). In the case of inclusions, anisotropy 226 can be observed in the effective medium to which the waves are sensitive if a degree of 227 alignment or preferred orientation persists over a broad area. This alignment of inclu-228 sions would result from local deformation, and thus also contain information about man-229 tle flow. However, studies observing high frequency scatterers in the lowermost mantle 230 observe very weak velocity contrasts (<0.1%: Mancinelli & Shearer, 2013). Extremely 231 strong isotropic velocity anomalies (10–30%) are only observed in thin patches of sev-232 eral 10s of km on top of the core-mantle boundary, the so-called ultra-low velocity zones 233 (e.g., Garnero et al., 1998; Yu & Garnero, 2018). 234

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### 3.1 Geodynamic models

Assumptions on the flow occurring in the lowermost mantle have varying degrees of complexities. In the simplest of models, horizontal flow is assumed causing simple horizontal shear as one might expect in a iso-chemical thermal boundary layer (as used in Wookey & Kendall, 2007). Anisotropy observed, however, does not have to represent local deformation, but could represent fossilized anisotropy. Anisotropic material can be

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formed elsewhere and be transported and rotated without overriding the preferred orientation. Therefore it is important to track the history of deformation for material in a given location. To represent change in flow direction in downwellings and upwellings in the lowermost mantle, corner flow streamlines can be used (Tommasi et al., 2018).

A range of studies use fully numerical models, where a number of assumptions on parameters for the lower mantle need to be made. The history of deformation is tracked by passive tracers that are advected through the model and record the velocity gradient at each step. The deformation history is generally used from the top of the lower mantle (e.g., Cottaar et al., 2014), or from the bridgmanite to post-perovskite transition (e.g Walker et al., 2011).

In one approach (Cottaar et al., 2014; Chandler et al., 2018), the CitcomS program (Zhong et al., 2000) is used to solve for the conservation of mass, momentum and energy, in a system that is heated from below, and where a slab is forced down from the top. Tracers are introduced at the top of the slab, and a large number of them eventually end up in the lowermost mantle, although the final distribution is irregular and shows clumping of tracers.

In a different approach (Walker et al., 2011; Nowacki et al., 2013), the flow field 261 is the instantaneous flow predicted by isotropic wave velocities, the gravity field, a 1D 262 viscosity model, and other geophysical constraints (Simmons et al., 2009). Because the 263 inversion assumes that flow does not change with time, regularly-spaced tracers can be 264 back-propagated to the top of D'' across the flow field, after which they are forward prop-265 agated to track the deformation along the path. The advantage of this method is that 266 one retrieves a regularly-sampled global anisotropic model that holds some potential re-267 lationship to the isotropic velocities, and that can thus be compared to global or regional 268 seismic observations. Additionally, this method tests a prior assumed relationship be-269 tween isotropic velocities and the flow field, testing models of thermal and/or thermo-270 chemical heterogeneity in the lower mantle. 271

While these geodynamical models represent test cases to explain lowermost seismic anisotropy, they are simplified in many ways. The geodynamical models have not explicitly included the bridgmanite to post-perovskite transitions, which only has a small density jump (Murakami et al., 2004; Oganov & Ono, 2004), but would cause significant viscosity weakening (Hunt et al., 2009) and allow slab material to spread more easily (Nak-

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agawa & Tackley, 2011). The viscosity model would be even more complex if the formation of CPO could be fed back into the geodynamical model creating anisotropic viscosity. So far, models have not tested the hypothesis of LLSVPs representing a different composition (e.g., Garnero et al., 2016), which appears important to understand the laterally varying anisotropy around LLSVP boundaries.

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### 3.2 Mineralogical constraints

In the upper mantle, the mineral olivine is abundant, and, with a highly anisotropic 283 crystal, represents a straightforward candidate to explain CPO anisotropy (Becker and 284 Lebedev, this volume). For the lowermost mantle the debate is still open as to which min-285 eral or polymineralic assemblage can explain the observed anisotropy. For a candidate 286 mineral or assemblage, we need to know its single crystal elasticity, which depending on 287 crystal symmetry can be described by three to 21 independent parameters. We mostly 288 rely on first-principle or *ab initio* calculations which solve the electronic Schrödinger equa-289 tion to obtain the crystal structure and the elasticity at high pressures and temperatures 290 (Buchen, this volume). Merely obtaining isotropic average elasticity information from 291 experiments under these extreme conditions is very challenging, let alone measuring the 292 independent anisotropic parameters (e.g., Marquardt et al., 2009; Finkelstein et al., 2018). 293

Additionally, we need to know how the candidate mineral or assemblage deforms 294 (Miyagi, this volume). To create seismic anisotropy, a mineral must significantly deform 295 by dislocation glide. In dislocation glide, dislocations within the crystal move along crys-296 tallographic planes. Preferred orientation results when crystals rotate to accommodate 297 glide along its weakest glide planes. Other mechanisms like diffusion creep or disloca-298 tion climb are not usually thought to cause preferred orientation, though this is not al-299 ways the case (Wheeler, 2009, 2010; Dobson et al., 2019). If dislocation glide is the pre-300 ferred mechanism, the next question that arises is what are the relative strengths of the 301 different slip systems (i.e., glide plane and slip directions). Calculations explore the rel-302 ative strengths of different deformation mechanisms and glide systems by calculating lat-303 tice friction and forces required to slip a dislocation (Peierls stress) in atomistic mod-304 els (Walker et al., 2010; Cordier et al., 2012). Experimentally, slip system activities can-305 not usually be measured for single crystals of the phases of interest here. Instead, ma-306 terials are deformed under compressive or shear stress in a large-volume apparatus (usu-307 ally on analogue materials), or in a diamond-anvil cell (For further details ?Romanow-308

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icz & Wenk, 2017) The resulting deformation may be imaged by X-ray diffraction. Dom inant slip systems may be estimated by inspection of the orientation distribution func tions (ODFs) of the crystallographic planes of interest, or inverted for by comparing for ward calculations of the experimental deformation with the results obtained.

To determine macroscopic anisotropy from these mineralogical constraints, the set of slip systems are combined with a deformation tensor to model a set of deformed crystals. Most often this is done using a homogenisation method such as the viscoplastic selfconsistent method (VPSC; Lebensohn & Tomé, 1993).

Forming the majority of the lowermost mantle, and thus the likeliest candidates to be the anisotropy-causing phases, are bridgmanite, post-perovskite and ferropericlase.

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### $3.2.1 \ Bridgmanite$

Bridgmanite, (Mg,Fe)SiO<sub>3</sub>-perovskite, is the most abundant mineral in the lower mantle (and in the Earth). Its pure Mg-endmember shows ~11% P wave and up to 15% S wave anisotropy (Oganov et al., 2001; Wentzcovitch et al., 2006; Stackhouse, Brodholt, Wookey, et al., 2005), and shows little (Li et al., 2005; Zhang et al., 2016) or variable (Fu et al., 2019) variation with the inclusion of iron.

There are mixed results on bridgmanite being a suitable candidate to explain anisotropy. Both experiments and calculations proposed a dominant glide plane of (001) (Wenk et al., 2004; Merkel et al., 2007; Ferré et al., 2007) which results in the opposite radial anisotropy to that observed (e.g., Wenk et al., 2011), while other experiments and calculations argue for a dominant (100) glide plane (Mainprice et al., 2008; Tsujino et al., 2016), which can create the observed  $V_{SH} > V_{SV}$  in simple shear. Miyagi & Wenk (2016) report a change from (001)-dominated glide to (100) around 55 GPa.

Bridgmanite is also known to be a very strong mineral. Experiments deforming a multi-phase mixture of bridgmanite and a smaller fraction of the weaker phase ferropericlase (or analogs), show in some cases that the ferropericlase takes up the majority of the deformation (Girard et al., 2016; Kaercher et al., 2016; Miyagi & Wenk, 2016), while in others the strong bridgmanite phase still dominates deformation (Wang et al., 2013) in line with simulations in a finite element model (Madi et al., 2005). Recently atomistic calculations have shown that the resistance to dislocation glide is very high, and dislocation climb should dominate (Boioli et al., 2017). Dislocation climb dominance could

- explain the general lack of anisotropy across most of the lower mantle, as well the high
- <sup>341</sup> viscosity of the lower mantle (Reali et al., 2019). However, attempts to explain weak anisotropy
- around ponded subducted slabs in the uppermost lower mantle in terms of bridgman-
- ite CPO (Tsujino et al., 2016; Walpole et al., 2017; Ferreira et al., 2019; Fu et al., 2019)
- would be therefore puzzling.
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### 3.2.2 Post-perovskite

Post-perovskite is a high-pressure polymorph of bridgmanite, which could become 346 stable in the lowermost mantle (Murakami et al., 2004; Oganov & Ono, 2004). Compared 347 to bridgmanite, post-perovskite is (1) more anisotropic (Iitaka et al., 2004; Stackhouse, 348 Brodholt, & Price, 2005; Wentzcovitch et al., 2006; Zhang et al., 2016), and (2) much 349 weaker to deform (Hunt et al., 2009; Ammann et al., 2010; Goryaeva et al., 2016). There-350 fore it is an attractive candidate to explain anisotropy observed in the lowermost man-351 tle. If, and to what degree, post-perovskite is actually stable at the pressures in the low-352 ermost mantle is still up for debate (see overviews in Cobden et al. (2015) and Hirose 353 et al. (2015)), but invoking the presence of post-perovskite helps explain S-to-P veloc-354 ity ratios in the lowermost mantle (P. Koelemeijer et al., 2018). If present, the strongly 355 positive Clapeyron slope of its phase transition from bridgmanite implies post-perovskite 356 is stable in a thicker layer in cold regions than in hot regions (Oganov & Ono, 2004; Tsuchiya 357 et al., 2004). Potentially post-perovskite becomes unstable again in the thermal bound-358 ary layer close to the core-mantle boundary, creating a lens of post-perovskite (Hernlund 359 et al., 2005). 360

Testing post-perovskite as a candidate to explain anisotropy is difficult as the pre-361 ferred slip system of post-perovskite is uncertain and diamond-anvil cell experimental 362 results have varied widely over the past 15 years (For further details, see ?, in this vol-363 ume.). The most recent results can be split in two categories. Experiments using  $MgSiO_3$ 364 post-perovskite, and  $MnGeO_3$  or  $MgGeO_3$  analogs, show a preferred slip plane of (001) 365 (e.g., Miyagi et al., 2010; Hirose et al., 2010; Nisr et al., 2012; X. Wu et al., 2017). Ex-366 periments using  $CaIrO_3$  postperovskite as an analog show a dominant slip system of [100](010) 367 (where [hkl] gives the Burgers vector; e.g., Yamazaki et al., 2006; Niwa et al., 2012; Hunt 368 et al., 2016). Atomistic models confirm the results of the latter category, showing both 369 slip systems [100](010) and [001](010) (Cordier et al., 2012; Goryaeva et al., 2015, 2017) 370

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as well as the occurrence of twinning  $1/2 < 110 > \{1\overline{1}0\}$  (Carrez et al., 2017). Addi-371 tionally, it is suggested that post-perovskite could inherit preferred orientation or tex-372 ture through the phase transition from bridgmanite (Dobson et al., 2013). Interpreta-373 tion of the texture inheritance in deformation experiments has been specifically argued 374 to explain part of the variation in interpreted preferred glide plane (e.g., Walte et al., 375 2009; Miyagi et al., 2011). This could only be the case if bridgmanite forms CPO tex-376 ture due to dislocation glide, which is debatable (Boioli et al., 2017). A more feasible sce-377 nario is bridgmanite inheriting texture from post-perovskite in a reverse transition which 378 could occur in the hotter regions (Dobson et al., 2013; Walker et al., 2018). 379

The importance of the incorporation of aluminium and iron into post-perovskite for our purposes depends on its effect on the stability field, deformation mechanisms, rheology and single-crystal anisotropy of the non-endmember phase. Iron- and aluminiumbearing post-perovskite is likely to be as anisotropic as the magnesian end-member at lowermost mantle conditions (??Zhang et al., 2016), but there is little evidence for its effect on plasticity. Recent work suggests that iron will strongly partition into ferropericlase in the lowermost mantle in any event (?), thus its importance may be limited.

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#### 3.2.3 Ferropericlase

Ferropericlase (Mg, Fe)O is present in the lower mantle with a molar abundance 388 of 10–30% (e.g., McDonough & Sun, 1995). Before post-perovskite was discovered in 2004, 389 ferropericlase was already considered a potential explanation of lowermost mantle anisotropy 390 (Yamazaki & Karato, 2002). It is cubic, and its elasticity can thus be described by three 391 independent parameters. These are constrained both through *ab initio* calculations (Karki 392 et al., 2000; Z. Wu et al., 2013) and through experiments (e.g., Jackson et al., 2006). The 393 results of these studies show significant single crystal anisotropy, as well as an increase 394 of anisotropy with Fe content, related to changes in  $C_{12}$  and  $C_{44}$  with cell volume via 395 pressure (Marquardt et al., 2009; Antonangeli et al., 2011; Finkelstein et al., 2018). 396

Ferropericlase is much weaker than bridgmanite (Cordier et al., 2012). Atomistic calculations of pure MgO endmember have shown dominating slip systems of  $\langle 110 \rangle \{1\bar{1}0\}$ and  $\langle 110 \rangle \{100\}$  (Carrez et al., 2009; Amodeo et al., 2011, 2016). Experiments on pure MgO (Merkel et al., 2002; Girard et al., 2012) and (Mg, Fe)O (Lin et al., 2009) show dominant slip on  $\langle 110 \rangle \{1\bar{1}0\}$ , while higher temperature experiments on (Mg, Fe)O also ac-

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tivate  $\langle 110 \rangle \{100\}$ , consistent with the calculations. Whether ferropericlase can explain 402 the observed anisotropy depends on the degree of single crystal anisotropy (related to 403 the Fe content), its abundance in the lowermost mantle (i.e. whether ferropericlase grains 404 become interconnected), and the general strength contrast between ferropericlase and 405 bridgmanite or post-perovskite. However, it should be noted that even in the two-phase 406 experiments discussed earlier, where ferropericlase takes up the bulk of the deformation, 407 coherent CPO does not develop in the ferropericlase, potentially due to the polyphase 408 geometry causing strain heterogeneity in the ferropericlase crystals (Kaercher et al., 2016; 409 Miyagi & Wenk, 2016). 410

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### 3.2.4 Other phases

412 Whilst post-perovskite, bridgmanite and ferropericlase are expected to dominate 413 the lowermost mantle, it is possible that other phases play a role in causing anisotropy.

Though peridotite comprises  $\sim 5\%$  of Ca-perovskite (CaMgSiO<sub>3</sub>) in the lower man-414 tle, basaltic compositions may hold up to 30% (McDonough & Sun, 1995), and hence 415 Ca-pv may be important if subducted material can accumulate at the base of the man-416 tle. Sample recovery issues mean that high-pressure and -temperature experiments are 417 difficult and the phase boundary between cubic and tetragonal Ca-pv is still being de-418 termined (Thomson et al., 2019), but molecular dynamics simulations (Li et al., 2006) 419 show maximum single-crystal shear wave anisotropy of 25%, similar to other phases men-420 tioned here. Room-temperature diamond-anvil cell experiments (Miyagi et al., 2009) and 421 Peierls–Nabarro modelling (Ferre et al., 2009) suggest Ca-pv might form a CPO in de-422 formation by glide on the cubic slip system  $\langle 1\bar{1}0\rangle \{110\}$ , and experiments on the analogue 423  $CaGeO_3$  suggest Ca-perovskite may be weaker than MgO (Wang et al., 2013), but rel-424 atively few studies have yet examined this further. 425

Silica phases may also make up  $\sim 20\%$  of a basaltic lower mantle. While it seems likely that stishovite is stable until about 1500 km depth, uncertainty remains as to when in the lower mantle silica transitions from the CaCl<sub>2</sub> structure to seifertite (e.g., Sun et al., 2019). This may be important since whilst seifertite appears to be only moderately anisotropic (?) and hence is likely not be a large contributor to lowermost mantle anisotropy, CaCl<sub>2</sub>-type silica may have much stronger shear wave anisotropy of about 30% (?). Un-

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fortunately we do not currently have constraints on how silica phases may accommodatestrain.

If hydrogen can be carried to the deep mantle, then hydrous phases such as aluminous phase D or phase H might occur in D" (e.g., ??), whilst aluminous phase  $\delta$ -AlOOH is likely present in basaltic compositions (e.g., ?), and iron-rich regions could contain Ferich phases such as FeO<sub>2</sub> or FeOOH (e.g., ?). Some of these phases may be strongly anisotropic, however compared to the nominally anhydrous silicates like bridgmanite and post-perovskite, little work has been done to understand their deformation mechanisms.

440

## 4 Joint geodynamic–seismic modelling

441

### 4.1 Recent developments

Several endeavours—mainly over the last decade—have tried to tie together all the fields and constraints discussed so far, in order to interpret anisotropy in the lowermost mantle. The long-term, sometimes enigmatic, goal of these studies is to map flow directions in the lowermost mantle to understand its role in the overall mantle convection (as the title of this chapter suggests). Most studies to this date, however, attempt to constrain the underlying cause of anisotropy taking their best guess at the flow regime.

In terms of the cause of anisotropy, recent studies rely heavily on post-perovskite being stable in the lowermost mantle to explain the observed anisotropy, arguing that bridgmanite produces the wrong radial anisotropic signature, and ferropericlase is not abundant enough to dominate anisotropic signatures. Bridgmanite not playing a major role can also be argued in light of the recent results that bridgmanite is too strong to cause dislocation glide and develop preferred orientation (Boioli et al., 2017).

454 The dominant glide plane in post-perovskite that is argued to explain anisotropy varies with studies arguing for dominant glide on (010) (Walker et al., 2011; Nowacki et 455 al., 2013; Creasy et al., 2017; Tommasi et al., 2018; Ford et al., 2015) and (001) (Nowacki 456 et al., 2010; Cottaar et al., 2014; Walker et al., 2018; X. Wu et al., 2017; Chandler et al., 457 2018). These studies range from finding a best fitting model from a qualitative compar-458 ison to previously published observations (i.e., Cottaar et al., 2014; Tommasi et al., 2018) 459 to a quantitative misfit with local observations (Nowacki et al., 2010; Ford et al., 2015; 460 Creasy et al., 2017) or with global anisotropic models (Walker et al., 2011, 2018). Of course, 461 all of these studies have made different assumptions and choices, which may affect the 462

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final conclusion. One example is the choice of elastic constants—for instance, a dominant glide system on the (010) plane results in  $V_{SH} > V_{SV}$  when using elastic constants of Stackhouse, Brodholt, & Price (2005), and in  $V_{SV} > V_{SH}$  when using those of Wentzcovitch et al. (2006) (Yamazaki & Karato, 2007; Wenk et al., 2011). Similarly, studies may choose to use constants derived at a single pressure and temperature—not necessarily those of the part of the mantle of interest—or attempt to include the variable anisotropic effects as P and T vary (Walker et al., 2011).

A number of recent studies are worth highlighting. Tommasi et al. (2018) explore the anisotropy resulting from deformation constraints from atomistic modelling instead of experimental results, arguing for a dominant glide plane of (010). Atomistic calculations, which will hopefully converge with experimental results in the future, offer a great step forward into constraining the deformation in the lowermost mantle. Their modelling finds weak radial anisotropy of  $V_{SH} > V_{SV}$  and sub-horizontal fast polarization directions in simple corner flow.

Their modelled elastic tensors with post-perovskite and periclase in an upwelling tracer can also fit the recent observations of changing anisotropy beneath Iceland by Wolf et al. (2019). However, the paper also presents models of pure bridgmanite and periclase that can fit the observations for the assumed change in flow.

Walker et al. (2018) explore texture inheritance from (001) slip in post-perovskite to bridgmanite (Dobson et al., 2013) on a global scale. Such a model can explain the observed sharp changes in the signature of anisotropy from regions dominated by cold downwellings, to regions dominated by hot upwellings or LLSVPs. Comparable results for texture inheritance were shown by Chandler et al. (2018) using the tracking of single tracers from downwelling to upwelling.

Most of the studies mentioned have pre-assumed the flow pattern either locally or globally and the models of different compositions are tested against seismic observations. Only the studies of Ford et al. (2015) and Creasy et al. (2017) both fit the compositional model as well as the flow direction. For both cases, this is applied to one locality where anisotropy is constrained from different azimuthal directions. Ford et al. (2015) suggest mainly vertical flows occur just to the East of the African LLSVP, while Creasy et al. (2017) suggest horizontal flows in a region of fast isotropic velocities beneath New Zealand

<sup>494</sup> and Australia.

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### 4.2 Example case: comparing a model to seismic observations

Many of the multi-disciplinary modelling studies discussed above compare their syn-496 thetic elasticity models to ray-theoretically derived local body wave observations. A num-497 ber of studies explore the limitations of interpreting body waves observations in terms 498 of anisotropy by analysing synthetic observations; for example for 1D isotropic or radi-499 ally anisotropic models for S<sub>diff</sub> waves (Maupin, 1994; Komatitsch et al., 2010; Borgeaud 500 et al., 2016; Parisi et al., 2018) and for ScS waves (Kawai & Geller, 2010). Nowacki & 501 Wookey (2016) extend the analysis for ScS waves to full synthetic anisotropic models with 502 small-scale variations from the model of Walker et al. (2011) (cf. Figure 3). They con-503 clude that ray-theoretical interpretations hold up for the simplest anisotropic models, 504 but break down for those with variable anisotropy. Additionally, the finite-frequency wave 505 senses less splitting than a ray-theoretical interpretation of an anisotropic model would 506 suggest, as the finite-frequency sensitivity will average over the strongly varying anisotropic 507 medium, sensing an effective medium. 508

Here we explore these limitations further by combining the forward modelling in a subducting slab of Cottaar et al. (2014) with the full-wave modelling of Nowacki & Wookey (2016) and analyse S<sub>diff</sub>, ScS, SKS, and SKKS phases in a finite-frequency framework.

512

### 4.2.1 Geodynamic and texture modelling

We use the results of Cottaar et al. (2014). We refer the reader to the original work for a full description of parameters used, but note that in this type of modelling, uncertainty can be introduced via a number of parameters including: the chosen relative critical resolved shear stresses on each slip system; the lack of a non-glide mechanism to accommodate strain and reset texture (such as diffusion creep); the phase boundary between phases; the single-crystal elastic constants; and the stress-strain homogenisation scheme. In the following we give a short summary of the modelling details.

Deformation is tracked along tracer particles (Figure 4a) across the lowermost mantle using CitcomS (Zhong et al., 2000), where 500 grains are modelled with the viscoplastic self-consistent method (VPSC; Lebensohn & Tomé, 1993) to accommodate the deformation. 75% of these grains are post-perovskite ('ppv') or bridgmanite ('pv'), and 25% are periclase. For post-perovskite, a dominant glide plane of (100) ('ppv 100' model), (010) ('ppv 010'), or (001) ('ppv 001') is assumed. The assumed glide planes for periclase are

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Table 1. Summary of synthetic paths used to investigate anisotropy in the geodynamic slab

544 model.

Code	Description, phases <sup>a</sup>	Source longitude (°)	Source latitude (°)	Focal mechanism <sup>b</sup>
А	Across slab: SKS–SKKS	0	-90	090/0/-90
В	Along slab: SKS–SKKS, $\mathbf{S}_{\mathrm{diff}}$	-55	0	180/30/0
$\mathbf{C}$	Along slab: ScS, $S_{diff}$	0	0	280/30/-90

<sup>a</sup> Phases analysed from full-wavefield synthetics.

<sup>b</sup> Given as strike/dip/rake of the fault plane in °.

All events at 650 km depth.

weaker, and this phase ends up accommodating 35-40% of the deformation. For the elastic constants, the values of Stackhouse, Brodholt, Wookey, et al. (2005) are used for postperovskite and bridgmanite, and those of Karki et al. (2000) for periclase. Cottaar et al. (2014) note that model ppv 001 is in general agreement with the radial anisotropy observed at the bottom of slabs and fast directions of azimuthal anisotropy is parallel to flow directions. The radial anisotropy in model ppv 010 also has the right sign, but is very weak in nature.

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### 4.2.2 Seismic modelling

We seek to compare the predicted seismic characteristics of our geodynamic slab 534 model to regional observations of anisotropy, including splitting in Sdiff, SKS-SKKS and 535 S–ScS differential splitting, as well as observations of changes in splitting intensity (Chevrot, 536 2000). In order to do this, we simulate the propagation of waves through the model in 537 two directions—along the slab and across it—for a range of geometries. We use synthet-538 ics in the epicentral distance range 55°  $\,\leq\,\,\Delta\,\,\leq\,\,80^\circ$  for ScS,  $100^\circ\,\,\leq\,\,\Delta\,\,\leq\,\,130^\circ$  for 539 SKS,  $110^{\circ} \le \Delta \le 130^{\circ}$  for SKKS and  $95^{\circ} \le \Delta \le 120^{\circ}$  for S<sub>diff</sub>. Table 1 outlines the 540 geometries used in this study and which phases are investigated for each, whilst Figure 4b 541 shows the location of the events and receivers. 542

We calculate the seismic response of the slab using the spectral element method as implemented in the SPECFEM3D\_GLOBE code (Komatitsch & Tromp, 2002). In or-

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der to remove time in writing intermediate files, we use a version of the code where creating the spectral element mesh and solving the equations of motion are performed in the same program (Komatitsch et al., 2003; Nowacki & Wookey, 2016). We use two chunks of the cubed sphere with 800 spectral elements along each side, giving seismograms accurate at frequencies below 0.2 Hz, similar to the dominant period of the waves at these distances.

The elasticity model is mapped to the seismic computational mesh by finding the 568 nearest neighbouring tracer particle within a defined 'slab' region, which is below 400 km 569 above the CMB and within 150 km from any given particle. Beyond this distance, the 570 nearest particle's elasticity grades smoothly to the background 1D velocity, given by AK135 571 (Kennett et al., 1995), over a 100 km distance using Voigt averaging between the isotropic 572 and full elastic tensor. This smoothing distance was chosen to avoid artificially extend-573 ing the region of the mantle influenced by the slab, whilst avoiding seismic artifacts from 574 a spatially abrupt transition between isotropic and anisotropic mantle. 575

576

#### 4.2.3 Synthetic analysis

We can process synthetics from our forward model in the same way as data and 577 compare the two. For the purposes of this example, we show a selection of results for the 578 three paths, for different combinations of seismic phases, analysing the shear wave split-579 ting in ScS, S<sub>diff</sub>, and differential splitting between SKS and SKKS. In all cases, we anal-580 yse the shear wave splitting in a window around the arrival of interest using the minimum-581 eigenvalue method of Silver & Chan (1991), with errors as updated by Walsh et al. (2013). 582 The fast axis here is defined as the angle  $\phi'$  from the radial component (or the vertical 583 at the bottoming point of the seismic ray, Figure 1c in Nowacki et al. (2010)). We also 584 consider the splitting intensity (SI; Chevrot, 2000) for SKS and SKKS waves, where the 585 polarisation is known to be radial, and S<sub>diff</sub> waves, where almost all SV energy is lost 586 along the diffracted path, rendering them horizontally polarised. For this reason,  $S_{diff}$ 587 SI is calculated in the opposite sense to usual for SK(K)S waves, interchanging the ra-588 dial and transverse components in the calculation. Discrepant SKS–SKKS splitting pairs 589 are identified where either one of the phases shows null splitting, whilst the other does 590 within error, or the two phases' 95% confidence region of the small-eigenvalue surface 591 do not overlap. Additionally, for all splitting measurements we use the automatic clas-592

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sification method of Wuestefeld et al. (2010) to calculate Q, a measure of quality between -1 and 1. -1 indicates a null, 1 a good measurement, and 0 a likely poor measurement.

<sup>595</sup> We consider first the splitting in ScS for path 'C' (Figure 5). This path is similar <sup>596</sup> to observations of splitting along palaeosubduction zones such as beneath the Caribbean <sup>597</sup> (Garnero et al., 2004; Maupin et al., 2005; Nowacki et al., 2010, e.g.,). In general, it seems <sup>598</sup> that  $\phi' \approx 90^{\circ}$  ( $\xi > 1$ ) in many slab regions (Nowacki et al., 2011; Romanowicz & Wenk, <sup>599</sup> 2017); in the Caribbean in particular, Garnero et al. (2004) infer a systematic rotation <sup>600</sup> in the fast angle across the palaeoslab region, giving a change in  $\phi'$  from ~ 105° to ~ <sup>601</sup> 75°.

Returning to Figure 5, it is clear that plasticity models ppv 010 and ppv 100 are 611 better candidates than the remaining models at reproducing the  $V_{\rm SH} > V_{\rm SV}$  and vari-612 able  $\phi'$  signals seen in data. Note that this is a different conclusion from Cottaar et al. 613 (2014), which could be due to the added complexity in this study of investigating the 614 non-horizontal orientation of the waves, which can cause rotations in  $\phi'$  (see Figure 2). 615 The models produce values of  $\delta t$  which are mostly comparable to those seen in nature, 616 though larger at up to 6 s in the synthetics versus  $\sim 2-3$  s as observed. The strength of 617 anisotropy present in the models is up to  $A^U = 0.1$ , which is about one-third of the pre-618 dicted single crystal anisotropy of ppv in the lowermost mantle (Stackhouse, Brodholt, 619 Wookey, et al., 2005). 620

We next show results for path 'B', which samples the slab similarly, but using  $S_{diff}$ and SKS–SKKS phases, in Figures 6 and 7. SI for  $S_{diff}$  should be large only when a significant non-radial anisotropy is present, which is generally the case within and at the edges of the slab.

It is notable in all cases that the pattern of  $\phi'$  and  $\delta t$  is complex and variable across 637 the slab, with large regions of null splitting even where strong anisotropy is present. Null 638 splitting may occur when the polarisation of a shear wave travelling through an anisotropic 639 medium is close to the fast or show shear wave orientations (which are perpendicular) 640 in that direction, and this may be the cause here. Variability in  $\phi'$  is expected because 641 of the pattern of flow in the model, and we observe fairly smooth rotations of  $\phi'$  from 642 north to south in the ppv models, similarly to data. However, the pv case shows  $\delta t$  vari-643 ations do not correlate strongly to simple features in the elasticity model. This illustrates 644 the sometime non-intuitive manner in which the seismic wave averages structure, and 645

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cautions against the use of approximate methods like ray theory when calculating syn-thetics in such models for data comparison.

SKS–SKKS pairs, in contrast, show relatively straightforward behaviour, with dis-648 crepant pairs concentrating near the edges of the high-anisotropy areas as expected. The 649 'core' of the deformed material does not show discrepant splitting, as both phases show 650 similar behaviour. Notably, for ppv 001 there is a region in the northeast which does not 651 show the expected behaviour. We see a similar non-intuitive behaviour in path A (Fig-652 ure 4b). Here, although some discrepant pairs straddle the edge of the slab for ppv 001, 653 very few paths show this for any of the other models. Inferring the edges of anisotropic 654 regions therefore must again be done with caution. 655

We also examine the raw difference in splitting intensity between SKS and SKKS 656 at the same seismograms,  $\Delta SI = SI_{SKS} - SI_{SKKS}$  (Figure 8). Measuring  $\Delta SI$  is com-657 putationally simple, and hence holds the promise for automatic global mapping of D''658 anisotropy. Comparing with the differential splitting predictions (Figures 4c, 6 and 7), 659 it appears that the along-slab path B (Figure 8b) shows straightforward behaviour, where 660  $\Delta$ SI deviates significantly from 0 where the slab is significantly anisotropic, either pos-661 itively or negatively depending on the exact elastic structure in the model. This agrees 662 well with the differential splitting interpretation. For path A, however (Figure 8a), a large 663 negative  $\Delta SI$  signal is present at the eastern end of most models, though this is not cor-664 related to significant differential splitting (Figure 4c). Most models also show large  $\Delta SI$ 665 in the central northern part, but again this is not reflected in differential splitting. This 666 suggests that although there is significant difference between the elastic structure expe-667 rienced by the SKS and SKKS waves in these regions as they cross the edge of the anisotropic 668 part of the slab, the shear wave splitting is not sufficiently coherent and clear to provide 669 a strong signal. Nevertheless, these calculations suggest that a more global SKS–SKKS 670 comparison holds promise for detecting regions where anisotropy changes rapidly. 671

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### 5 Limitations, advances, and the way forward

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#### 5.1 The inverse problem

The final goal of observing anisotropy in the lowermost mantle is, as the title of this chapter suggests, to map flow directions. We have discussed the forward model and the large number of assumptions required to create an anisotropic model and compare

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it to seismic observations. For most studies discussed in Section 4.1 the flow model was
 one of the prior assumptions and different potential compositions are explored; only the
 recent studies of Ford et al. (2015) and Creasy et al. (2017) locally interpret flow direc tion.

Results for a suite of candidate deformation mechanisms, like those in Figure 5, im-685 mediately make tempting a potentially circular line of reasoning: given an assumed flow 686 model and the data, can we infer the deformation mechanism responsible for anisotropy? 687 And with that improved estimate of deformation mechanism, can we then infer the flow 688 field? As discussed, uncertainty as to the very cause of anisotropy in the lowermost man-689 the makes such reasoning perilous. It is also worth noting that if the rheology in geody-690 namic models is set in part based on observations of seismic anisotropy, and the assump-691 tion made of a particular deformation mechanism, then there is an added danger in the 692 use of such dynamic models to then infer the mechanism of anisotropy. 693

Despite these problems, we can proceed with caution if we hold in mind that it is the *combination* of the flow model, deformation mechanism and mineral elasticity which is being tested against the data in each instance, not any one of these in isolation. Intuitively, varying any one of these might lead to an equally well-fitting set of synthetic observations when varying another.

Is finding the dominant mineral (or multi-phase system) and deformation systems creating lowermost mantle anisotropy the biggest hurdle in the way to mapping flow? If we constrain the main source of anisotropy, could we create a map of flow across the mantle? To do that, we would be interested in making a number of inverse steps (shown in Figure 3 by the white dashed arrows), each of which is non-linear and under-determined:

1. Using the seismic observations to find the constrained parts of the seismic anisotropic 704 tensor. Constraints on the anisotropic tensor will always be limited by the prop-705 agation direction of the seismic phases used (Figure 2) and the azimuthal cover-706 age, generally leaving large parts of the anisotropic tensor unconstrained, and re-707 verting studies to assume symmetries (i.e. radial or azimuthal anisotropy). Ad-708 ditionally, the resulting anisotropic tensor would always reflect the effective elas-709 tic tensor that the seismic waves observed at long wavelengths, and could result 710 from an entire suite of small-scaled heterogeneously (an)isotropic media (e.g., Backus, 711 1962; Capdeville & Cance, 2014; Fichtner et al., 2013). In this chapter we have 712

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714 715

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frequently made the major assumption that the anisotropy of the effective tensor is due to underlying intrinsic anisotropy (CPO), and not caused by small-scale isotropic heterogeneity (SPO).

2. Mapping from the effective anisotropic tensor to a set of textured minerals or a 716 specific mineral with preferred orientation. Accounting for the null-space in the 717 elastic tensor, there would be no unique fit here and many assumptions on the min-718 eral physics need to be made. One could not account for the entire suite of po-719 tential deformation mechanisms occurring, presence of other minerals and the re-720 lated multi-phase deformation effects. With the assumption of a single dominant 721 mineral and glide mechanism, the main imaged fast polarisation direction in the 722 elastic tensor would be preferentially fit. There is no value in mapping the strength 723 of the anisotropy into a degree of preferentially aligned minerals, as the amplitudes 724 of the effective elastic tensor will be underestimated (as shown by the synthetic 725 results in Section 4.2). 726

3. Mapping from textured minerals to potential deformation history and flow directions. A single textured mineral can result from various deformation histories as
it can both reflect the deformation it is undergoing, or fossilised deformation, which
could be displaced and rotated. To uniquely constrain the flow model, many observations in different locations will have to be combined, as well as including other
constraints, i.e. isotropic velocities, the gravity field, and past plate tectonic models.

While we choose to highlight the true inverse steps of this problem, this poses mul-734 tiple layers of non-uniqueness, which makes a flow map for the lowermost mantle based 735 on anisotropy appear unobtainable. For the foreseeable future, mapping mantle flow will 736 have to rely on simplified relationships between fast polarisation directions and flow with 737 understanding of the conditions under which these are valid. For the upper mantle, a 738 simple relationship is posed as the fast axis of deformed olivine generally aligns with the 739 flow direction, which has allowed interpretation of asthenospheric flow from anisotropy, 740 although the validity of under volatile-rich conditions, e.g. in the mantle wedge above 741 subduction zones, where observations also become more complex (see Becker and Lebe-742 dev, this volume). For the lowermost mantle, Cottaar et al. (2014) pose a similar rela-743 tionship between horizontal fast direction and horizontal flow direction specifically for 744 post-perovskite with dominant (001)-glide for a simple case of a slab spreading out on 745

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the CMB. Such a relationship should be tested statistically under many flow conditions,
rotations, and deformations, and using full-waveform modelling. Tommasi et al. (2018)
present a relationship of sub-parallel fast polarisation directions to the flow direction for
post-perovskite with dominant (010)-glide and twinning for simple horizontal flow. Their
study tests, not statistically but very systematically, the limitations of this relationship
in corner flows and the ability of different seismic phases to detect these polarisation directions.

Creasy et al. (2019) pose a different question: with our current level of non-uniqueness in the interpretation, how many independent seismic observations do we need in a single location to interpret composition and flow? They show statistically that roughly 10 or more measurements of fast direction or reflection polarisations with various azimuths and incidence angles are needed to uniquely constrain the anisotropic tensor to make an interpretation with some confidence. It is challenging to find locations to which this can be applied due to available earthquake-station geometries.

Whilst it is unlikely that the mineralogical parameters we have discussed will be 760 tightly constrained for some time, and similarly seismic data coverage will probably not 761 improve vastly, it is conceivable that probabilistic approaches to inferring flow from anisotropy 762 may enable progress by incorporating uncertainties in all the input parameters as in Fig-763 ure 3 and retrieving an ensemble of acceptable flow histories. Such a model suite would 764 however likely be a vast undertaking, requiring many millions of forward iterations, in-765 cluding geodynamic and full waveform modelling. This is unfeasible with the current com-766 bination of forward numerical methods and computational resources, but the future may 767 bring this within our grasp. 768

#### 5.2 Outlook

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While mapping flow clearly remains an ambitious goal, current studies of anisotropy do provide new insights into the deepest mantle. Specifically, the goal to find the source of anisotropy reveals the potential importance of post-perovskite to be stable in the mantle, as bridgmanite might be too strong to cause texturing (Boioli et al., 2017) and ferropericlase, as the minor phase, might not deform coherently (Miyagi & Wenk, 2016). Observations of lateral changes in anisotropy could highlight where post-perovskite is present, which relates to the temperature field and thus the convective patterns. A ma-

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jor step forward in understanding the role of post-perovskite would be to resolve its dominant deformation mechanisms. It remains to be seen if the latest theoretical calculations
(e.g., Goryaeva et al., 2017) will converge with future experimental results.

Studies of anisotropy are also illuminating the nature of the LLSVP boundaries. 780 From a seismological point of view, the claim that the radial anisotropy switches sign 781 inside and outside the LLSVPs needs to be further tested for its robustness. Any rela-782 tionship between the isotropic and anisotropic velocities in tomographic models could 783 be an artefact (e.g., Chang et al., 2015). Local observations of splitting, however, have 784 confirmed strong changes in anisotropy around the edges of the LLSVPs (e.g., Cottaar 785 & Romanowicz, 2013; Wang & Wen, 2007; Lynner & Long, 2014). The nature of the LLSVPs 786 poses major unanswered questions, and understanding the changing signatures of anisotropy 787 can help resolve to what degree their boundaries represent a purely thermal or a thermo-788 chemical gradient. In the thermal case, change in anisotropy could be explained by a phase 789 transition from post-perovskite to perovskite (Dobson et al., 2013) or by a change in flow 790 direction, likely from horizontal outside to vertical within the LLSVPs or plumes (e.g. 791 Wolf et al., 2019). In the case where LLSVPs represent thermo-chemical piles, the bound-792 ary could also be mechanical with separate convection inside and outside the piles (e.g., 793 Garnero & McNamara, 2008). Currently, capturing all the variation in parameters which 794 contribute to the development of anisotropy whilst correctly relating these to observa-795 tions is computationally constrained. However, while we are far from producing a global 796 flow map based on anisotropic variations, anisotropic studies play a role in answering 797 these fundamental questions on the nature of the lowermost mantle. With the answers 798 to these questions, flow can be more easily interpreted on the basis of mapped isotropic 799 velocity variations. 800

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Figure 1. a. -d. Maps show shear wave radial anisotropic parameter  $\xi$  at 2800 km depth 95 for SEMUCBwm1 (French and Romanowicz, 2014), S362ANI+M (Moulik and Ekstrom 2014), 96 SAVANI (Auer et al. 2014), SGLOBE-rani (Chang et al., 2015). Red and blue contours show 97 bounds at three votes for the isotropic slow and fast cluster based on votes across five isotropic 98 models (Cottaar & Lekić, 2016). Histograms show distribution and mean values of  $\xi$  in the differ-99 ent cluster vote areas, red-'slow', blue-'fast', yellow-'neutral' (boundaries for 'neutral' cluster are 100 not shown on the maps). e. Vote map showing where models agree on  $\xi > 1$  or  $\xi < 1$  ( $V_{SH}$  or 101  $V_{SV}$  being faster, respectively). All models agree that  $\xi~>~1$  for 18% of the map, while for  $\xi~<~1$ 102 the area is 13%. 103



Figure 2. a. Ray paths for general body wave (pairs) used to constrain lowermost mantle 201 anisotropy: SKS–SKKS (blue), ScS–S (red), S<sub>diff</sub> (green) (made with Obspy; Krischer et al., 202 2015). b.-e. Hemisphere projections of various assumed anisotropic symmetries viewed from 203 above (made with MSAT; Walker & Wookey, 2012). Bars show splitting direction and bar 204 lengths show splitting strength as a function of shear wave propagation direction. Coloured shad-205 ing shows general sensitivity of body waves (see a.) although there is some overlap. b. Radial 206 anisotropy with  $\xi = 1.03$  c. 3% azimuthal anisotropy with a fast axis direction of 112°. d. Tilted 207 anisotropy, i.e. anisotropy in c. tilted by  $40^\circ.$ e. Full anisotropic tensor for 75% post-perovskite 208 and 25% periclase in a downgoing slab (see Section 4.2). 209



Figure 3. Showing the general steps in forward modelling from a flow model to interpreting seismic observations following the filled blue arrows as discussed in Section 3. Dashed arrows indicate the inverse steps to go from seismic observations to flow directions, for which challenges and limitations are discussed in Section 5



(a) Geodynamical setup of forward calculation, taken from Figure 1 of Cottaar et Figure 4. 545 al. (2014). A slab is imposed, moving from north to south, which subducts along the equator. 546 Tracer particles are shown by orange dots, with the path for two highlighted as green and white 547 lines. Section a-a' along longitude  $0^{\circ}$  to the right shows non-dimensional temperature. (b) Ge-548 ometry of synthetic seismic sources and receivers in relation to the slab model. Sources are shown 549 by colour-coded lower-hemisphere focal mechanisms (annotated with the code in Table 1), match-550 ing the receiver locations, shown by open areas with solid boundaries. Red lines show the paths 551 of tracer particles, and orange dash-dotted lines show the slab edges in panel (a). Areas shown 552 by other figures are indicated by dotted black lines and labelled. (c) Discrepant SKS-SKKS split-553 ting for path 'A' (Table 1) for each plasticity model. Red and orange circles respectively show 554 the core piercing points for SKKS and SKS waves for pairs which are discrepant, whilst white 555 circles denote the piercing points of pairs which are not. Underlying colour shows the strength of 556 anisotropy at the bottom of the slab texture model, using the universal elastic anisotropy index, 557  $A^{U}$  (Ranganathan & Ostoja-Starzewski, 2008) according to the colour scale on the bottom right. 558 (For approximate path lengths of SKS and SKKS in the lowermost mantle, see Figure 6.) 559



Shear-wave splitting results for the ScS phase for path 'C' (Table 1), for each Figure 5. 602 slab texture. Each panel shows on the left the fast shear wave orientation in the ray frame (see 603 Nowacki et al., 2010, Figure 1c),  $\phi'$ , as a black bar located at the ScS core bounce point. Bars 604 oriented left–right ( $\phi' = 90^{\circ}$ ) correspond to radial anisotropy with  $\xi > 1$ , and vertical bars mean 605  $\xi < 1 \ (\phi' = 0^{\circ})$ , with non-radial anisotropy otherwise (when  $0^{\circ} \neq \phi' \neq 90^{\circ}$ ). Colour beneath the 606 bars is as in Figure 4c. On the right we show the amount of splitting,  $\delta t$  at each bounce point, 607 coloured by the second scale bar. Crosses signify null measurements. The length of the ray path 608 of ScS in the lowermost 250 km of the mantle is shown in the right hand panel for the ppv 001 609 case. 610



Figure 6. Shear-wave splitting results for the path 'B' (Table 1) and models ppv001 and 625 ppv010. For each plasticity model, four panels show: (top left) the splitting intensity (Chevrot, 626 2000) of  $S_{diff}$  as colour; (bottom left) the ray-frame fast shear wave orientation of  $S_{diff}$ , coloured 627 by the splitting quality measure Q (Wuestefeld et al., 2010); (bottom right) slow shear-wave 628 delay time  $\delta t$ , coloured as per the scale bar; and (top right) pairs of discrepant SKS-SKKS 629 splitting, where the red and orange circles show the core piercing point of SKKS and SKS, re-630 spectively, and white circles indicate no discrepant splitting; background colour shows strength 631 of anisotropy as in Figure 4c. S<sub>diff</sub> points are plotted at the end of the core-diffracted part of 632 the path. The lengths of the ray paths of  $S_{diff}$ , SKS and SKKS in the lowermost 250 km of the 633 mantle are shown in the SI panel for the ppv 001 case. 634



Figure 7. Shear-wave splitting results for the path 'B' (Table 1) and models ppv100 and pv.
Features as for Figure 6.



Figure 8. Difference in splitting intensity between SKS and SKKS,  $\Delta$ SI, for (a) path A and (b) path B (Table 1).  $\Delta$ SI is shown by colour according to the scale, lower right, at the midpoint between SKS and SKKS core-mantle boundary piercing points. Grey lines in the background show the path of tracers particles in the geodynamic model.