SPARC CCMVal PhotoComp-2008 (version 1 July 2008)

GOALS. Evaluate how models calculate photolysis (and indirectly heating) rates in the stratosphere and troposphere with the incentive of locating errors or biases and identifying improved and practical methods. There are three basic parts to PhotoComp2008:

(1) Basic test of *all* J- values for high sun (SZA=15°), w/ & w/o additional scattering layers (stratiform clouds & stratospheric volcanic aerosols).

(2) Test of twilight, sphericity, and 24-hour averages (SZA = $84^{\circ} - 96^{\circ}$).

(3) Test of wavelength integration w/o scattering (SZA = 15°).

There will be one standard atmosphere, whose primary definition will include air mass, ozone mass, and temperature in each layer. This atmosphere is typical of the tropics, ozone column = 260 DU. For efficiency, we will use this same atmosphere in all sections, even the low-sun, polar cases.

PARTICIPATION. This study is designed to aid development and testing of the photolysis and short-wave heating codes used in chemistry-transport models and coupled chemistry-climate models. This project is open: any research group can participate by running the experiments and reporting the results as specified below. We also encourage participation from groups (without CTMs or CCMs) who have participated in other model-measurement studies (e.g., IPMMI, POLARIS). Many CTM/CCMs will be using "the same" photolysis scheme (e.g., fast-TUV, fast-J) and think their participation redundant – this is false. The implementation of a standard scheme into any CTM/CCM will likely alter (intended or inadvertent) how the J-values are calculated: thus it is very important when you perform these tests that the photolysis module that is as close a possible to that embedded within the CTM/CCM and not the original, standalone version that you used to derive your inline model.

EXPERIMENTS.

Part 1 is a basic test of all J-values for high sun (SZA = 15°) over the ocean (albedo = 0.10, Lambertian). **Part 1a: Clear sky** (only Rayleigh scattering) and no aerosols. **Part 1b: Pinatubo aerosol** in the stratosphere (layer 10). **Part 1c: Stratus cloud** (layer 2). The primary atmosphere (Table 1a) is specified in terms of pressure layers, mean temperature, and column O_3 in each layer. Please do not include absorption by NO₂ or other species in calculating optical depths. For 1b and 1c we recommend that you use the specified optical properties in Table 1c, interpolating across the 5 specified wavelengths.

Part 2 tests the simulation of a spherical atmosphere and twilight conditions that are critical to the polar regions. Use the same atmosphere as Part 1 without clouds or aerosols. Assume equinox (solar declination = 0°) and a latitude of 84°N. The surface SZA (not including refraction) varies from 84° (noon) to 96° (midnight). Report all J-values at noon, midnight, and the 24-hour average (integrating as you would in your CTM/CCM). With a spherical atmosphere, the local solar zenith angle changes with altitude and if refraction is included it will change the surface angle. Please note how you treat the solar ray path in your model description.

Part 3 tests the accuracy of wavelength binning in the critical region 290-400 nm that dominates tropospheric photolysis. Shut off all Rayleigh scattering and surface reflection (albedo = 0) giving effectively a simple Beer's Law calculation. Repeat the calculation in Part 1, but report only J-values for J-O3 (i.e., total), J-O3(1d) $[O_3 => O_2 + O(^1D)]$, and J-NO2 $[NO_2 => NO + O]$. These are the two critical J-values for the troposphere, and they both have unusual structures in absorption cross section and quantum yields. The organizers will make these calculations using very high resolution (0.05 nm) cross sections and solar fluxes and for different options (e.g., JPL-06 vs. IUPAC cross sections) to provide a benchmark. NOTE that we will only use results below 20 km (L=1:11) for this comparison.

DIAGNOSTICS.

Model Documentation should include a brief outline of the methods and any references (limit: one page). Please include brief notes on: how you treat sphericity and refraction, the Schumann-Runge bands (J-O₂ and J-NO), Rayleigh scattering, multiple scattering, clouds and aerosols, seasonal changes in sun-earth distance, solar variability, and any specific parameterizations. Default cross sections are JPL-2006, please note if you are using alternate.

Report all J-values and all standard model layers since this is a check on all modeled Jvalues, not just the radiative transfer solution. See Appendix for data formatting. We are not specifying the day-of-the-year, so use solar fluxes for sun-earth distance = 1.0 au and average over the 11-yr solar cycle if possible. UCI's high-resolution solar spectrum used in these experiments is the average of two high and low SUSIM spectra (29 Mar 1992) and 11 Nov 1994), this is not meant to be the 11-yr average. It will be provided at 0.05 nm resolution, but we encourage you to use your own solar fluxes for the primary tests since changing solar fluxes will mostly likely require a complete re-averaging of all cross-sections (see Fast-J paper, Wild, Zhu, Prather, 2000). Please report in model documentation what you are using for the solar spectrum and how the solar cycle is represented in your submissions, and if possible submit it as a separate file so that it may be used to address differences later. (With different wavelength binning, this will not be trivial.) Reported photolysis rates should be calculated for the mass mid-point of each layer, this brings PhotoComp closer to current CTM usage rather than the original gridpoint formulation used in M&M. Results in the form of clearly labeled ascii text files should be uploaded to BADC CCMVal archive or emailed to the organisers (see web posting for specific details).

DISCUSSION.

Implementation into a particular model's code will up to the participant. For example, at UCI we have two models that we will use in PhotoComp: a fast-JX model within the CTM that uses layers of uniform composition defined by mass (kg/m^2) ; and a stand-alone photochemical box model that defines altitude (in cm) as the vertical grid and uses number densities for air and ozone. For the latter, we have re-mapped the primary

atmosphere (Table 1a) onto a grid-point structure (Table 1b) that has the same mid-layer properties as the layer mean value and the same columns of O_2 and O_3 .

One question will be: What is the correct answer? In some cases we may be able to define a "best" answer based on obvious physics or convergence of some of the more resolved models, but in others we may not. Thus in all of our proposed experiments we will begin with a "standard model" result (not necessarily the best answer) from one of the models and then determine a best answer, if possible, after analysis of the results.

One approach to defining the correct answer would be to merge observed radiation fields or photolysis rates (e.g., IPMMI, POLARIS, see references below), but we feel this may be too difficult to match the exact observing conditions. One way to include the knowledge gained by these field studies is to ensure participation from some of the models (e.g., NCAR-TUV, APL).

We do not recommend reporting detailed actinic fluxes as a function of wavelength since everyone selects different ways of integrating over wavelength (e.g., bins) and trying to reconcile the different wavelength scales is not worthwhile. If major problems show up, then a subgroup of models can consider how to resolve the differences.

Another major issue with photolysis and heating rates is the treatment of clouds and cloud fraction. This is very important, but probably beyond the current PhotoComp. It would require a special workshop. We do include an option for a plane-parallel volcanic aerosol layer (aka Pinatubo) and a stratiform cloud.

APPENDIX

Standard Atmosphere & Other Specifications

Table 1a. PhotoComp 2008 standard atmosphere L edge p(hPa) T 03(mass/mass) DU(**redund. 1 1000.0 299.9 3.844E-08 2.4532 2^^ 866.0 289.5 4.704E-08 4.8514 3 649.4 278.8 4.720E-08 2.8831 5 365.2 253.9 5.551E-08 2.4140 6 273.8 239.7 5.977E-08 1.9491 7 205.4 224.6 6.390E-08 1.6527 8 154.0 209.4 9.012E-08 1.6527 9 115.5 198.2 1.486E-07 2.0441 10** 86.60 195.8 3.885E-07 4.0065 11 64.94 203.1 1.533E-06 11.8582 12 48.70 209.9 3.790E-06 29.7855 13 36.52 215.5 6.849E-06 29.7855 14 27.38 220.1 1.034E-05 33.7361 15	
$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	08 standard atmosphere
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	T 03(mass/mass) DU(**redundant)
16 15.40 228.9 1.577E-05 28.9124 17 11.55 233.3 1.653E-05 22.7293 18 8.660 237.8 1.670E-05 17.2239 19 6.494 242.6 1.578E-05 12.2053 20 4.870 248.3 1.363E-05 7.9054 21 3.652 254.1 1.236E-05 5.3734 22 2.738 259.5 9.733E-06 3.1740 23 2.054 262.9 8.158E-06 1.9951 24 1.540 265.1 6.721E-06 1.2325	9.9 3. 844E -08 2. 4532 9.5 4. 704E -08 4. 8514 8.8 4. 720E -08 3. 6498 7.2 4. 972E -08 2. 8831 3.9 5. 551E -08 2. 4140 9.7 5. 977E -08 1. 9491 4.6 6. 390E -08 1. 6527 9.4 9. 012E -08 1. 6527 8.2 1. 486E -07 2. 0441 5.8 3. 885E -07 4. 0065 3.1 1. 533E -06 11. 8582 9.9 3. 790E -06 21. 9783 5.5 6. 849E -06 29. 7855 0.1 1. 034E -05 33. 7361 4.5 1. 326E -05 32. 4219 8.9 1. 577E -05 28. 9124 3.3 1. 653E -05 22. 7293 7.8 1. 670E -05 17. 2239 2.6 1. 578E -05 12. 2053 8.3 1. 363E -05 7. 9054 4.1 1. 236E -05 5. 3734 9.5 9. 733E -06 3. 1740 2.9 8. 158E -06 1. 9951
25 1.155 266.9 5.511E-06 0.7578 26 0.8660 264.7 4.810E-06 0.4960 27 0.6494 261.8 4.009E-06 0.3100 28 0.4870 259.7 3.325E-06 0.1928	6.9 5.511E-06 0.7578 4.7 4.810E-06 0.4960 1.8 4.009E-06 0.3100 9.7 3.325E-06 0.1928

	29	0.3652	254.3	2.820E-06	0. 1226		
	30	0. 2738	247.0	2.323E-06	0. 0758		
	31	0.2054	239.4	1.909E-06	0. 0467		
	32	0. 1540	234.4	1.585E-06	0. 0291		
	33	0. 1155	232.7	1.335E-06	0. 0184		
	34	0.08660	226.4	1. 102E-06	0. 0114		
	35	0.06494	216.4	8. 927E-07	0. 0069		
	36	0.04870	210. 8	7.372E-07	0.0043		
	37	0.03652	208.0	6. 168E-07	0.0027		
	38	0.02738	205.2	5.162E-07	0.0017		
	39	0.02054	202.4	4.317E-07	0.0011		
	40	0.01540	199.4	3.607E-07	0.00066		
	41	0.01155	197.6	3.032E-07	0.00042		
	. 42	0.00866					
ć	above m	nodel top lay	yer=41 @ (D. 0866 hPa,	assume uniform 1 & 03		
	42	0.00866	197.6	3.032E-07	0.00125		
	43	0.00000					
** Lover for Stratiform cloud: UD (600 nm) = 20.0							
and Tayer for Prhatubo Surfate aerosol: UD(600 nm) = 1.00							
see Table TC.							

Table 1b. Standard atmosphere shown mapped into grid points

Q = scattering efficiency (average of cross-section / (pi $\,$ * r**2)) typically Q $\,$ 2 for large clouds and large aerosols

K = extinction (m2/g), the cross-sectional area per gram of material K(m2/g) = 0 / [4/3 * Reff(micron) * Rho(g/cm3)]Rho = density of particles (g/cm3) n = index of refraction OD = optical depth (column) = column mass (g/m2) * K (m2/g)SSA = single scattering abledo $LG(1:8) = coefficients of Legendre expansion of scattering phase fn. \\ both polariztions are added. By definition SLEG(1) = 1. \\ Fast-JX uses these first 8 terms to define the scattering.$ g = asymmetry factor = LG(2) / 3. Pinatubo: OD = 1.0 in layer 10 (86.6 to 64.9 hPa) _____ ======== Stratospheric aerosol composed of 75%-wt H2SO4. Rho = 1.630= 1.514 + 0.000i (200 nm) n 1.473 + 0.000i (300 nm) 1.459 + 0.000i (400 nm) 1. 448 + 0. 000i 1. 435 + 0. 000i (600 nm) (999 nm) Log-normal distribution with RO = 0.08 micron & sigma = 0.800 **check that you are using the right log-normal by deriving Reff Reff = 0.386 micron K (600nm) = 2.610 OD (@600nm) = 1.00 ==> aerosol = 1.00/K = 0.3832 g/m2 W Q SSA LG(2) LG(3) LG(4) LG(5) LG(6) LG(7) LG(8) 200 2.5935 1.0000 2.092 2.914 2.880 3.295 3.185 3.430 3.379 300 2.6669 1.0000 2.121 2.861 2.792 2.936 2.733 2.703 2.568 400 2.5588 1.0000 2.144 2.813 2.711 2.695 2.425 2.257 2.069 600 2.1893 1.0000 2.149 2.713 2.547 2.362 2.018 1.740 1.499 999 1.4540 1.0000 2.118 2.537 2.277 1.951 1.555 1.229 0.972 (fast-JX v61 scatter #15) Stratus: OD = 20.0 in layer 2 (866 to 649 hPa) _____ Pure water cloud Rho = 1.000 n = 1.335 + 0.000i (assumed 200-999 nm) Deirmendjian Cumulus C1 (Gamma, $n(r) = a r^*alpha exp[-b r^*gamma])$ mode radius Rc = 4 microns, al pha=6, b=3/2, gamma = 1 Reff = 6.00 micron K (600nm) = 0.2668OD (@600nm) = 20.0 ==> aerosol = 20.0/K = 75.0 g/m2 W Q SSA LG(2) LG(3) LG(4) LG(5) LG(6) LG(7) LG(8) 200 2.0650 1.0000 2.610 3.998 4.771 5.450 6.196 6.829 7.721 300 2.0835 1.0000 2.596 3.973 4.725 5.406 6.129 6.751 7.607 400 2.1064 1.0000 2.571 3.936 4.660 5.345 6.056 6.670 7.492 600 2.1345 1.0000 2.557 3.902 4.596 5.263 5.923 6.507 7.267 999 2.1922 1.0000 2.499 3.799 4.418 5.081 5.667 6.213 6.851 (fort 40.45) (fast-JX v61 scatter #08) ____ _____
 Table 2.
 Standard diagnostics and file names
 -----Ascii tables will be fine given small data sets. Report J-values at the mid-point of Layers 1 through 40. File names: PCO8_{model name + version if need be}_{PhotoPart#} Write format: J-title, J-value(1:41) '(a8, 1x, 41e9. 2)' File Examples: PC08_UCI ref_doc. txt (or .pdf or .doc if need formatting) UCI old reference code, documentation PC08_UCI-JX_doc. txt UCI version of fast-JX, documentation PC08_UCI ref_P1a. txt

fn of aerosol size distrib N(r), index of refraction, wavelength.

Table 3. Standard J-value names.

Please use these abbreviations (if possible in the following order) so that J's can be sorted. For new J's please add with unique name. (available as PC08_J-labels.txt) Note that for some J's, the branching ratios do not have different cross-sections associated with them and the branching ratios are fixed, hence we report only one J. For many organics, the quantum yields are complex and have been incorporated into these J's. If you do not calculate one of these, please keep that row in your table with zero or blank values.

_____ J-NO NO =N+0 2 J-02 02 =0+0 =0+02 (total = both 0(3P) and 0(1D))* =0(1D)+02 3 J-03 03 J-03(1d) J-H2C0a 4 03 H2COa 5 =H+HCÓ 6 J-H2C0b H2C0b =H2+C0 7 J-H202 H202 =OH+OH 8 J-CH300H CH300H =CH30+0H J-N02 9 N02 =N0+0 =N0+02(11.4%) & N02+0(88.6%)* 10 J-N03 N03 J-N205 =N02+N03 11 N205 J-HN02 HONO =OH+NO 12 13 J-HNO3 HNO3 =0H+N02 14 J-HNO4 H02N02 = 0H + N03J-CI NO3a CINO3a =CI+NO3 15 CI N03b =CI 0+N02 CI 2 =CI +CI J-CI NO3b 16 J-CI 2 17 HOCI J-HOCI =OH+CI18 J-0CI 0 19 0CI 0 =0+CI 0 CI 202 =CI +CI +O2 J-CI 202 20 21 J-CIOCLO =CI + 022 J-Br0 Br0 =Br+0 J-BrN03 BrN03 =Br+N03(29%) & Br0+N02(71%)* 23 24 J-HOBr HOBr =0H+Br 25 J-BrCI BrCl =Br+Cl 26 J-N20 N20 =N2+0 CFCI 3 =... 27 J-CFCI 3 CF2CI2 =28 J-CF2CI 2 CF2CI CFCI 2=. . . J-F113 29 30 J-F114 CF2CI CF2CI = . . . 31 J-F115 CF3CF2CI = . . . CCI 4 =. CH3CI =C 32 J-CCI 4 33 J-CH3CI =CH3+CI CH3CCI 3=... 34 J-MeCCL3 35 J-CH2CI 2 CH2CI 2 = . . . 36 J-CHF2CI CHF2CI =... 37 J-F123 CF3CHCI 2=. . . 38 J-F141b CH3CFCI 2=. . . J-F142b 39 CH3CF2CI =. CH3Br =CH3+Br J-CH3Br 40 J-H1211 CF2CI Br=... 41 CF3Br =... 42 J-H1301 43 J-H2402 C2F4Br2=... 44 J-CH2Br2 CH2Br2 =... CHBr3 =. J-CHBr3 45 CH3I =CH3+I CF3I =CF3+I J-CH3I 46 47 J-CF3I

J-OCS J-PAN OCS =C0+S CH3C(0)02N02 =CH3C(0)02+N02(60%) & CH3C(0)0+N03(40%)* CH30N02=CH30+N02 48 49 J-CH3N03 50 CH3CH02=CH3CHC0 CH3C(0)=CH3+HC0 CH3C(0)CH=CH2 =C3H6+C0(60%) & CH2=CHC0+CH3(40%)* CH2C(CH3)CH0 =CH2=C(CH3)+HC0 H0CH2CH0 =H0CH2+HC0 CH3C0C2H5 =CH3+C2H5C0(15%) & C2H5+CH3C0(85%)* 51 52 J-ActAl d J-MeVK 53 54 55 J-MeAcr J-GI yAI d J-MEKeto C2H5CHO =C2H5+HCO CH3COCHO =CH3CO+HCO 56 J-EAI d
 Sol J-EAI O
 CC2/SCHO
 =CE/SCHO

 57
 J-MGI yxI
 CH3COCHO
 =CH3CO+HCO

 58
 J-GI yxI a
 (CHO)2
 =HCO+HCO

 59
 J-GI yxI b
 (CHO)2
 =H2+CO+CO

 61
 J-Acet-a
 C3H6O
 =CH3CO+CH3

 62
 J-Acet-b
 C3H6O
 =CH3+CH3+CO

* In preliminary comparisons, we have found it best to compare the total 03 photolysis rate and the rate leading to O(1D), skipping the O(3P) path. When branching paths with % are indicated in the table, they indicate the values derived for fast-JX, please just report the total J-value.